

Essentials of Geology

TWELFTH EDITION

Frederick K. Lutgens • Edward J. Tarbuck • Illustrated by Dennis Tasa

ALWAYS LEARNING



ESSENTIALS OF GEOLOGY

12e GLOBAL EDITION

Frederick K. Lutgens Edward J. Tarbuck

ILLUSTRATED BY

Dennis Tasa

PEARSON

Boston Columbus Indianapolis New York San Francisco Upper Saddle River Amsterdam Cape Town Dubai London Madrid Milan Munich Paris Montréal Toronto Delhi Mexico City São Paulo Sydney Hong Kong Seoul Singapore Taipei Tokyo

Acquisitions Editor: Andrew Dunaway	Editorial Assistant: Sarah Shefveland
Head of Learning Asset Acquisition, Global Editions: Laura Dent	Senior Marketing Assistant: Nicola Houston
Acquisition Editor, Global Editions: Priyanka Ahuja	Full Service/Composition: Cenveo® Publisher Services
Assistant Project Editor, Global Editions: Paromita Banerjee	Project Manager, Full Service: Heidi Allgair
Senior Marketing Manager: Maureen McLaughlin	Photo Manager: Maya Melenchuk
Project Manager: Crissy Dudonis	Photo Researcher: Kristin Piljay
Project Management Team Lead: Gina M. Cheselka	Text Permissions Manager: Alison Bruckner
Senior Production Manufacturing Controller, Global Editions: Trudy Kimber	Design Manager: Derek Bacchus
Executive Development Editor: Jonathan Cheney	Interior Design: Elise Lansdon Design
Director of Development: Jennifer Hart	Cover Designer: Lumina Datamatics
Content Producer: Timothy Hainley	Photo and Illustration Support: International Mapping
Project Manager, Instructor Media: Eddie Lee	Operations Specialist: Christy Hall
Media Production Manager, Global Editions: Vikram Kumar	Cover Image Credit: $\ensuremath{\mathbb O}$ Im Perfect Lazybones/Shutterstock

Credits and acknowledgments borrowed from other sources and reproduced, with permission, in this textbook appear on the appropriate page within text or are listed below.

Page 24: Quote from Aristotle, translated by Adams, F.D., in The Birth and Development of the Geological Sciences, Dover Publications, 1954; Page 25: Quote from James Hutton, Theory of the Earth, 1785; Page 25: Quote from William L. Stokes, Essentials of Earth History, Prentice Hall, Inc. 1973, p. 20; Page 26: Quote from James Hutton, Transactions of the Royal Society of Edinburgh, 1788; Page 28: Quote from Jacob Bronowski, The Common Sense of Science, p. 148, Harvard University Press, 1953; Page 29: Quote from F. James Rutherford and Andrew Ahlgren, Science for All Americans (New York: Oxford University Press, 1990), p. 7; Page 29: Quote from Speech delivered at Douai on December 7, 1854 on the occasion of his formal inauguration to the Faculty of Letters of Douai and the Faculty of Sciences of Lille), reprinted in: Pasteur Vallery-Radot, ed., Oeuvres de Pasteur (Paris, France: Masson and Co., 1939), vol. 7, page 131; Page 55: Quote from Alfred Wegener, The Origin of Continents and Oceans, translated from the 4th revised German ed. of 1929 by J. Birman (London: Methuen, 1966); Page 56: R. T. Chamberlain, quoted from Hallam, A. (1973) A Revolution in the Earth Sciences. Clarendon Press, Oxford; Page 124: Quote from Lee Green, MD, an associate professor at the University of Michigan Medical School;

Earth Sciences, Clarendon Press, O. MD, an associate professor at the U *Pearson Education Limited* Edinburgh Gate Harlow Essex CM20 2JE

England

and Associated Companies throughout the world

Visit us on the World Wide Web at: www.pearsonglobaleditions.com

© Pearson Education Limited 2015

The rights of Frederick K. Lutgens and Edward J. Tarbuck to be identified as the authors of this work have been asserted by them in accordance with the Copyright, Designs and Patents Act 1988.

Authorized adaptation from the United States edition, entitled Essentials of Geology, 12th Edition, ISBN 978-0-321-94773-4 by Frederick K. Lutgens and Edward J. Tarbuck, published by Pearson Education © 2015.

All rights reserved. No part of this publication may be reproduced, stored in a retrieval system, or transmitted in any form or by any means, electronic, mechanical, photocopying, recording or otherwise, without either the prior written permission of the publisher or a license permitting restricted copying in the United Kingdom issued by the Copyright Licensing Agency Ltd, Saffron House, 6–10 Kirby Street, London EC1N 8TS.

All trademarks used herein are the property of their respective owners. The use of any trademark in this text does not vest in the author or publisher any trademark ownership rights in such trademarks, nor does the use of such trademarks imply any affiliation with or endorsement of this book by such owners.

ISBN 10: 1-292-05718-1 ISBN 13: 978-1-292-05718-7

British Library Cataloguing-in-Publication Data

A catalogue record for this book is available from the British Library.

 $10\ 9\ 8\ 7\ 6\ 5\ 4\ 3\ 2\ 1$

Typeset in Minion Pro by Laserwords Private, Ltd Printed and bound by Neografia in Slovakia

Page 188: Quote from Jack Eddy, "A Fragile Seam of Dark Blue Light," in Proceedings of the Global Change Research Forum. U.S. Geological Survey Circular 1086, 1993, p. 15; Page 277: Quote from Walter Mooney, a USGS seismologist; Page 371: Quote from Exploration of the Colorado River of the West (Washington, DC: Smithsonian Institution, 1875), p. 203; Page 435: Quote from J. D. Hays, John Imbrie, and N. J. Shackelton, "Variations in the Earth's Orbit: Pacemaker of the Ice Ages," Science 194 (1976): 1121-32. p. 1131; Page 453: Quote from R. A. Bagnold, The Physics of Blown Sand and Desert Dunes, 2005; Page 494: Quote from James Hutton, Transactions of the Royal Society of Edinburgh, 1805; Page 504: Quote from B. Bryson, A Short History of Nearly Everything (Broadway Books, 2003); Page 563: Quote from IPCC, "Summary for Policy Makers." In Climate Change 2013: The Physical Science Basis; Page 569: Quote from J. T. Overpeck, et al., "Arctic System on Trajectory to New, Seasonally Ice-Free States," EOS, Transactions, American Geophysical Union, 86 (34): 309, August 23, 2005; Page 571: National Assessment Synthesis Team, Climate Change Impacts on the United States: The Potential Consequences of Climate Variability and Change (Washington, DC: U.S. Global Research Program, 2000), p. 19.

ind SmartFigures and Mobile Field Trip Figures

In addition to the many informative and colorful illustrations and photos throughout this text, you will find two kinds of special figures that offer additional learning opportunities. These figures contain QR codes which the student can scan with a smart phone to explore exciting expanded online learning materials.

where you see this icon.

Find **SmartFigures** where you see this icon.

Find Mobile Field Trip Figures

Chapter 1

- 1.5 Earth history—Written in the rocks (p. 25)
- Magnitude of geologic time (p. 27) 1.7
- Nebular theory (p. 35) 1.16
- Earth's layers (p. 38) 1.18
- The rock cycle (p. 41) 1.21
- 1.23 The continents (p. 44)

Chapter 2

- 2.2 Reconstructions of Pangaea (p. 53)
- Rigid lithosphere overlies the weak asthenosphere (p. 58) 2.9
- Continental rifting: Formation of a new 2 14 ocean basin (p. 61)
- 2.16 Three types of convergent plate boundaries (p. 62)
- 2.20 Transform plate boundaries (p. 66)
- Movement along the San Andreas Fault (p. 67) 2.22
- 2.30 Time scale of magnetic reversals (p. 73)

Chapter 3

- 3.3 Most rocks are aggregates of minerals (p. 87)
- Color variations in minerals (p. 92) 3.12
- 3.15 Common crystal habits (p. 93)
- Hardness scales (p. 94) 3.16
- 3.18 Cleavage directions exhibited by minerals (p. 95)
- 3.24 Five basic silicate structures (p. 98)

Chapter 4

- 4.3 Intrusive versus extrusive igneous rocks (p. 115)
- 4.7 Igneous rock textures (p. 119)
- 4.12 Classification of igneous rocks (p. 122)
- 4.24 Partial melting (p. 131)
- Sill exposed in Sinbad country, Utah (p. 133) 4.27

Chapter 5

- 5.10 Anatomy of a volcano (p. 153)
- 5.14 Cinder cone (p. 156)
- 5.23 Super-eruptions at Yellowstone (p. 165)
- Earth's zones of volcanism (p. 170) 5.29
- 5.30 Subduction of the Juan de Fuca plate produced the Cascade volcanoes (p. 172)
- 5.31 Large basalt provinces (p. 172)

Chapter 6

- Ice breaks rock (p. 182) 6.4
- 6.5 Unloading leads to sheeting (p. 182)

- The formation of rounded boulders (p. 186) 6.10
- Rock type influences weathering (p. 186) 6.11
- 6.13 Monuments to weathering (p. 187)

Chapter 7

- 7.2 The big picture (p. 205)
- Sorting and particle shape (p. 208) 7.7
- 7.17 Bonneville salt flats (p. 214)
- 7.22 Utah's Capitol Reef National Park (p. 219)
- 7.24 Lateral change (p. 222)
- 7.32 Common oil traps (p. 227)

Chapter 8

- 8.3 Sources of heat for metamorphism (p. 238)
- 8.4 Confining pressure and differential stress (p. 238)
- 8.14 Gneiss with garnet porphyroblasts, Adironacks, New York (p. 245)
- 8.19 Rocks produced by contact metamorphism (p. 248)
- 8.23 Metamorphism along a fault zone (p. 250)
- 8.24 Textural variations caused by regional metamorphism (p. 251)

Chapter 9

- 9.5 Elastic rebound (p. 260)
- 9.10 Body waves (P and S waves) versus surface waves (p. 263)
- Turnagain Heights slide caused by the 1964 Alaska Earth-9.26 quake (p. 271)
- 9.34 Seismac gaps: Tools for forecasting earthquakes (p. 277)

Chapter 10

- 10.4 Satellite altimeter (p. 289)
- 10.8 Active continental margins (p. 292)
- 10.17 Ridge segments that exhibit fast, intermediate, and slow spreading rates (p. 298)
- 10.22 Midcontinent rift (p. 302)
- 10.24 The demise of the Farallon plate (p. 302)

Chapter 11

- 11.1 Deformed sedimentary strata (p. 310)
- 11.6 Common types of folds (p. 314)
- Sheep Mountain, Wyoming (p. 315) 11.7
- 11.8 Domes versus basins (p. 315)
- 11.15 Normal dip-slip fault (p. 317)
- 11.16 Normal faulting in the Basin and Range Province (p. 318)
- 11.26 Collision and accretion of small crustal fragments to a continental margin (p. 325)
- 11.29 India's continued northward migration severely deformed much of China and Southeast Asia (p. 327)
- 11.30 Formation of the Appalachian Mountains (p. 328)
- 11.31 The Valley and Ridge Province (p. 329)
- 11.33 The effects of isostatic adjustment and erosion on mountainous topography (p. 330)

Chapter 12

- 12.2 Excavating the Grand Canyon (p. 339)
- 12.15 Gros Ventre rockslide (p. 347)
- 12.19 Creep (p. 350)
- 12.21 When permafrost thaws (p. 351)

Chapter 13

- 13.2 The hydrologic cycle (p. 359)
- 13.4 Mississippi River drainage basin (p. 360)
- 13.9 The Mighty Missippi near Memphis, Tennesee (p. 363)
- 13.12 Channel changes from head to mouth (p. 365)
- 13.17 Formation of cut banks and point bars (p. 369)
- 13.25 Incised meanders (p. 373)

Chapter 14

- 14.5 Water beneath Earth's surface (p. 389)
- 14.12 Hypothetical groundwater flow system (p. 393)
- 14.14 Cone of depression (p. 395)
- 14.16 Artesian systems (p. 396)
- 14.22 How a geyser works (p. 399)

Chapter 15

- 15.6 Movement of a glacier (p. 418)
- 15.9 Zones of a glacier (p. 419)
- 15.14 Erosional landforms created by alpine glaciers (p. 422)
- 15.21 Formation of a medial moraine (p. 426)
- 15.24 Common depositional landforms (p. 428)
- 15.34 Orbital variations (p. 435)

Chapter 16

- 16.1 Dry climates (p. 443)
- 16.8 Landscape evolution in the Basin and Range region (p. 447)
- 16.14 Formation of desert pavement (p. 451)

- 16.16 White Sands National Monument (p. 453)
- 16.17 Cross-bedding (p. 453)
- 16.19 Types of sand dunes (p. 454)

Chapter 17

- 17.6 Passage of a wave (p. 464)
- 17.10 Wave refraction (p. 466)
- 17.11 The longshore transport system (p. 467)
- 17.15 Some depositional features (p. 470)
- 17.25 East coast estuaries (p. 476)
- 17.29 Hurricane source regions and paths (p. 478)

Chapter 18

- 18.7 Inclusions (p. 495)
- 18.8 Formation of an angular unconformity (p. 493)
- 18.13 Applying principles of relative dating (p. 496)
- 18.18 Fossil assemblage (p. 500)
- 18.21 Radioactive decay curve (p. 502)

Chapter 19

- 19.4 Major events that led to the formation of early Earth (p. 518)
- 19.10 Growth of continents (p. 523)
- 19.12 The major geologic provinces of North America (p. 524)
- 19.15 Connection between ocean circulation and the climate in Antarctica (p. 526)
- 19.28 Relationships of vertebrate groups and their divergence from lobe-finned fish (p. 535)

Chapter 20

- 20.5 Ice cores: Important sources of climate data (p. 551)
- 20.16 Paths taken by solar radiation (p. 557)
- 20.18 The greenhouse effect (p. 558)
- 20.24 Monthly CO₂ concentrations (p. 562)
- 20.34 Slope of the shoreline (p. 570)

Brief Contents

- 1 An Introduction to Geology 20
- 2 Plate Tectonics: A Scientific Revolution Unfolds 50
- 3 Matter and Minerals 84
- **4** Igneous Rocks and Intrusive Activity 112
- 5 Volcanoes and Volcanic Hazards 144
- 6 Weathering and Soils 178
- 7 Sedimentary Rocks 202
- 8 Metamorphism and Metamorphic Rocks 234
- 9 Earthquakes and Earth's Interior 256
- **10** Origin and Evolution of the Ocean Floor 286
- **11** Crustal Deformation and Mountain Building 308
- **12** Mass Wasting: The Work of Gravity 336
- **13** Running Water 356
- 14 Groundwater 384
- **15** Glaciers and Glaciation 412
- 16 Deserts and Wind 440
- 17 Shorelines 458
- **18** Geologic Time 488
- **19** Earth's Evolution Through Geologic Time 512
- 20 Global Climate Change 546

Appendix

Metric and English Units Compared 576

Glossary 577

Index 586

Contents

PREFACE 15

1 An Introduction to Geology 20

- **Geology: The Science of Earth 22** Physical and Historical Geology 22
 Geology, People, and the Environment 23
- 1.2 The Development of Geology 24 Catastrophism 24 The Birth of Modern Geology 25 Geology Today 25 The Magnitude of Geologic Time 26
- 1.3 The Nature of Scientific Inquiry 27

Hypothesis Theory **28** Scientific Methods Plate Tectonics and Scientific Inquiry

1.4 Earth's Spheres 30 Hydrosphere 30 Atmosphere 31 Biosphere 32

Geosphere 32

- 1.5 Earth as a System 33 Earth System Science 33 The Earth System 34
- 1.6 Early Evolution of Earth 35 Origin of Planet Earth 35 Formation of Earth's Layered Structure 36

Earth's Internal Structure 37
 Earth's Crust 37
 Earth's Mantle 37
 Earth's Core 39

1.8 Rocks and the Rock Cycle 39 The Basic Cycle 39 Alternative Paths 40

1.9 The Face of Earth 42Major Features of the Continents **44**Major Features of the Ocean Floor **45**

Concepts in Review 46 Give It Some Thought 49

2 Plate Tectonics: A Scientific Revolution Unfolds 50

- 2.1 From Continental Drift to Plate Tectonics 52
- 2.2 Continental Drift: An Idea Before Its Time 53
 Evidence: The Continental Jigsaw Puzzle 53
 Evidence: Fossils Matching Across the Seas 54
 Evidence: Rock Types and Geologic Features 54
 Evidence: Ancient Climates 55
- 2.3 The Great Debate 56 Rejection of the Drift Hypothesis 56
- 2.4 The Theory of Plate Tectonics 57
 Rigid Lithosphere Overlies Weak Asthenosphere 58
 Earth's Major Plates 58
 Plate Boundaries 58
- 2.5 Divergent Plate Boundaries and Seafloor Spreading 60
 Oceanic Ridges and Seafloor Spreading 60
 Continental Rifting 61
- 2.6 Convergent Plate Boundaries and Subduction 62

 Oceanic–Continental Convergence 63
 Oceanic–Oceanic Convergence 64
 Continental–Continental Convergence 65
- 2.7 Transform Plate Boundaries 66
- 2.8 How Do Plates and Plate Boundaries Change? 68 The Breakup of Pangaea 69 Plate Tectonics in the Future 69
- 2.9 Testing the Plate Tectonics Model 70
 Evidence: Ocean Drilling 70
 Evidence: Mantle Plumes and Hot Spots 71
 Evidence: Paleomagnetism 72
- 2.10 How Is Plate Motion Measured? 75 Geologic Evidence for Plate Motion 75 Measuring Plate Motion from Space 76

2.11 What Drives Plate Motions? 77

Forces That Drive Plate Motion Models of Plate–Mantle Convection Concepts in Review Give It Some Thought

3 Matter and Minerals 84

- 3.1 Minerals: Building Blocks of Rock 86 Defining a Mineral 86 What Is a Rock? 87
- 3.2 Atoms: Building Blocks of Minerals 88 Properties of Protons, Neutrons, and Electrons 88 Elements: Defined by Their Number of Protons 88

3.3 Why Atoms Bond 90

The Octet Rule and Chemical Bonds Ionic Bonds: Electrons Transferred Covalent Bonds: Electron Sharing Metallic Bonds: Electrons Free to Move

3.4 Properties of a Mineral 92 Optical Properties 92

> Mineral Strength **93** Density and Specific Gravity **95** Other Properties of Minerals **95**

3.5 Mineral Groups 96 Classifying Minerals 96

Silicate Versus Nonsilicate Minerals **96**

3.6 The Silicates 97

Silicate Structures **97** Joining Silicate Structures **98**

- 3.7 Common Silicate Minerals 99 The Light Silicates 99 The Dark Silicates 101
- 3.8 Important Nonsilicate Minerals 102
- 3.9 Minerals: A Nonrenewable Resource 105 Renewable Versus Nonrenewable Resources 105 Mineral Resources and Ore Deposits 105

Concepts in Review: 107 Give It Some Thought 110

4 Igneous Rocks and Intrusive Activity 112

- 4.1 Magma: Parent Material of Igneous Rock 114 The Nature of Magma 114 From Magma to Crystalline Rock 115 Igneous Processes 115
- 4.2 Igneous Compositions 116

 Granitic (Felsic) Versus Basaltic (Mafic)
 Compositions 116
 Other Compositional Groups 117
 Silica Content as an Indicator of Composition 117
- 4.3 Igneous Textures: What Can They Tell Us? 118 Types of Igneous Textures 118
- 4.4 Naming Igneous Rocks 121

Granitic (Felsic) Igneous Rocks Andesitic (Intermediate) Igneous Rocks Basaltic (Mafic) Igneous Rocks Pyroclastic Rocks

4.5 Origin of Magma 125 Generating Magma from Solid Rock 125

- 4.6 How Magmas Evolve 127
 Bowen's Reaction Series and the Composition of Igneous Rocks 127

 Magmatic Differentiation and Crystal Settling 128
 Assimilation and Magma Mixing 129
- 4.7 Partial Melting and Magma Composition 130 Formation of Basaltic Magma 130 Formation of Andesitic and Granitic Magmas 130
- 4.8 Intrusive Igneous Activity 131

Nature of Intrusive Bodies Tabular Intrusive Bodies: Dikes and Sills Massive Intrusive Bodies: Batholiths, Stocks, and Laccoliths

4.9 Mineral Resources and Igneous Processes 135
 Magmatic Segregation and Ore Deposits 135
 Hydrothermal Deposits 137
 Origin of Diamonds 137

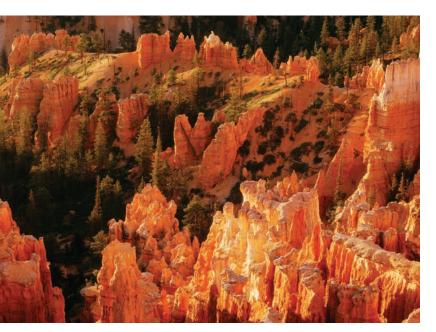
 Concepts in Review 138

Give It Some Thought 142

5 Volcanoes and Volcanic Hazards 144

- 5.1 Mount St. Helens Versus Kilauea 146
- 5.2 The Nature of Volcanic Eruptions 147 Factors Affecting Viscosity 147 Quiescent Versus Explosive Eruptions 148





- 5.3 Materials Extruded During an Eruption 149
 Lava Flows 149
 Gases 151
 Pyroclastic Materials 151
- 5.4 Anatomy of a Volcano 153

5.5 Shield Volcanoes 154 Mauna Loa: Earth's Largest Shield Volcano 154

Kilauea, Hawaii: Eruption of a Shield Volcano **155** Cinder Cones **156**

5.6 Cinder Cones 156 Parícutin: Life of a Garden-Variety Cinder Cone 157

5.7 Composite Volcanoes 158

5.8 Volcanic Hazards 159

Pyroclastic Flow: A Deadly Force of Nature **159** Lahars: Mudflows on Active and Inactive Cones **160** Other Volcanic Hazards **161**

5.9 Other Volcanic Landforms 163 Calderas 163 Fissure Eruptions and Basalt Plateaus 165

Lava Domes **166** Volcanic Necks and Pipes **166**

5.10 Plate Tectonics and Volcanic Activity 168 Volcanism at Convergent Plate Boundaries 169 Volcanism at Divergent Plate Boundaries 169 Intraplate Volcanism 172

Concepts in Review 173 Give It Some Thought 176

6 Weathering and Soils 178

- 6.1 Weathering 180
- 6.2 Mechanical Weathering 181

Frost Wedging Salt Crystal Growth Sheeting **183** Biological Activity

- 6.3 Chemical Weathering 184
 Water and Carbonic Acid 184
 How Granite Weathers 184
 Weathering of Silicate Minerals 185
 Spheroidal Weathering 185
- 6.4 Rates of Weathering 186 Rock Characteristics 186 Climate 187 Differential Weathering 187
- 6.5 Soil 188 An Interface in the Earth System 188 What Is Soil? 188
- 6.6 Controls of Soil Formation 189
 Parent Material 189
 Time 189
 Climate 189
 Plants and Animals 190
- Topography 190 6.7 The Soil Profile 191
- 6.8 Classifying Soils 192
- 6.9 The Impact of Human Activities on Soil 194

Clearing the Tropical Rain Forest—A Case Study of Human Impact on Soil **194** Soil Erosion: Losing a Vital Resource **195**

6.10 Weathering and Ore Deposits 198 Bauxite 198 Other Deposits 198

> Concepts in Review 198 Give It Some Thought 201

7 Sedimentary Rocks 202

- 7.1 The Importance of Sedimentary Rocks 204
- 7.2 Origins of Sedimentary Rock 205
- 7.3 Detrital Sedimentary Rocks 206 Shale 207 Sandstone 208 Conglomerate and Breccia 210
- 7.4 Chemical Sedimentary Rocks 210 Limestone 210 Dolostone 212 Chert 213 Evaporites 214

- 7.5 Coal: An Organic Sedimentary Rock 214
- 7.6 Turning Sediment into Sedimentary Rock: Diagenesis and Lithification 216
 Diagenesis 216
 Lithification 216
- 7.7 Classification of Sedimentary Rocks 217
- 7.8 Sedimentary Rocks Represent Past Environments 218 Importance of Sedimentary Environments 218 Sedimentary Facies 219

Sedimentary Structures 222

- 7.9 Resources from Sedimentary Rocks 224 Nonmetallic Mineral Resources 224 Energy Resources 225
- 7.10 The Carbon Cycle and Sedimentary Rocks 227 Concepts in Review 229 Give It Some Thought 232

8 Metamorphism and Metamorphic Rocks 234

- 8.1 What Is Metamorphism? 236
- 8.2 What Drives Metamorphism? 237 Heat as a Metamorphic Agent 237 Confining Pressure 238 Differential Stress 239 Chemically Active Fluids 239 The Importance of Parent Rock 240
- 8.3 Metamorphic Textures 240 Foliation 240 Foliated Textures 241 Other Metamorphic Textures 242
- 8.4 Common Metamorphic Rocks 243 Foliated Metamorphic Rocks 243 Nonfoliated Metamorphic Rocks 245
- 8.5 Metamorphic Environments 246 Contact or Thermal Metamorphism 246 Hydrothermal Metamorphism 247 Burial and Subduction Zone Metamorphism 248 Regional Metamorphism 248 Other Metamorphic Environments 249

8.6 Metamorphic Zones 250 Textural Variations 250 Index Minerals and Metamorphic Grade 250 Concepts in Review 252

Give It Some Thought 254

9 Earthquakes and Earth's Interior 256

- 9.1 What Is an Earthquake? 258
 Discovering the Causes of Earthquakes 259
 Aftershocks and Foreshocks 260
 Faults and Large Earthquakes 261
- 9.2 Seismology: The Study of Earthquake Waves 262 Instruments That Record Earthquakes 262 Seismic Waves 263
- 9.3 Locating the Source of an Earthquake 264
- 9.4 Determining the Size of Earthquakes 266 Intensity Scales 266 Magnitude Scales 266

9.5 Earthquake Destruction 269 Destruction from Seismic Vibrations 269 Landslides and Ground Subsidence 270 Fire 270 What Is a Tsunami? 271

- 9.6 Where Do Most Earthquakes Occur? 273
 Earthquakes Associated with Plate boundaries 273
 Damaging Earthquakes East of the Rockies 274
- 9.7 Can Earthquakes Be Predicted? 275 Short-Range Predictions 275 Long-Range Forecasts 277
- 9.8 Earth's Interior 278
 Formation of Earth's Layered Structure 279
 Probing Earth's Interior: "Seeing" Seismic Waves 279



9.9

Earth's Layers 280 Crust 280 Mantle 280 Core 281 Concepts in Review 281

Give It Some Thought 284

10 Origin and Evolution of the Ocean Floor 286

- 10.1 An Emerging Picture of the Ocean Floor 288 Mapping the Seafloor 288 Provinces of the Ocean Floor 290
- 10.2 Continental Margins 290 Passive Continental Margins 290 Active Continental Margins 291
- 10.3 Features of Deep-Ocean Basins 292
 Deep-Ocean Trenches 292
 Abyssal Plains 293
 Volcanic Structures on the Ocean Floor 293
 Explaining Coral Atolls—Darwin's Hypothesis 294
- 10.4 Anatomy of the Oceanic Ridge 295
- 10.5 Oceanic Ridges and Seafloor Spreading 296 Seafloor Spreading 297 Why Are Oceanic Ridges Elevated? 297 Spreading Rates and Ridge Topography 297
- 10.6 The Nature of Oceanic Crust 298 How Does Oceanic Crust Form? 299 Interactions Between Seawater and Oceanic Crust 299
- 10.7 Continental Rifting: The Birth of a New Ocean Basin 300 Evolution of an Ocean Basin 300



Failed Rifts 302

 10.8 Destruction of Oceanic Lithosphere 302 Why Oceanic Lithosphere Subducts 302 Subducting Plates: The Demise of Ocean Basins 303 Concepts In Review 304 Give It Some Thought 307

11 Crustal Deformation and Mountain Building 308

- 11.1 Crustal Deformation 310 What Causes Rocks to Deform? 311 Types of Deformation 311 Factors That Affect Rock Strength 312
- 11.2 Folds: Rock Structures Formed by Ductile Deformation 313
 Anticlines and Synclines 313
 Domes and Basins 314
 Monoclines 315
- 11.3 Faults and Joints: Rock Structures Formed by Brittle Deformation 316
 Dip-Slip Faults 317
 Strike-Slip Faults 318
 Joints 319
- 11.4 Mountain Building 321
- 11.5 Subduction and Mountain Building 322
 Island Arc–Type Mountain Building 322
 Andean-Type Mountain Building 322
 Sierra Nevada, Coast Ranges, and Great Valley 323
- 11.6 Collisional Mountain Belts 324

 Cordilleran-Type Mountain Building 324
 Alpine-Type Mountain Building: Continental Collisions 326
 The Himalayas 326
 The Appalachians 327
- 11.7 What Causes Earth's Varied Topography? 330
 The Principle of Isostasy 330
 How High Is Too High? 331

 Concepts in Review 332
 Give It Some Thought 334

12 Mass Wasting: The Work of Gravity 336

12.1 The Importance of Mass Wasting 338
 Landslides as Geologic Hazards 338
 The Role of Mass Wasting in Landform
 Development 339

Slopes Change Through Time 339

- 12.2 Controls and Triggers of Mass Wasting 340 The Role of Water 340 Oversteepened Slopes 341 Removal of Vegetation 341 Earthquakes as Triggers 342 Landslides Without Triggers? 343 The Potential for Landslides 343
- 12.3 Classification of Mass-Wasting Processes 344
 Type of Material 344
 Type of Motion 344
 Rate of Movement 345
- 12.4 Rapid Forms of Mass Wasting 345 Rockslide 346 Debris Flow 347 Earthflow 348
- 12.5 Slow Movements 349 Creep 349 Solifluction 350 The Sensitive Permafrost Landscape 350 Concepts in Review 351 Give It Some Thought 354

13 Running Water 356

- 13.1 Earth as a System: The Hydrologic Cycle 358
- 13.2 Running Water 359 Drainage Basins 360 River Systems 360 Drainage Patterns 361
- 13.3 Streamflow 362 Factors Affecting Flow Velocity 362 Changes Downstream 364
- 13.4 The Work of Running Water 365 Stream Erosion 365 Transport of Sediment by Streams 366 Deposition of Sediment by Streams 368
- 13.5 Stream Channels 368 Bedrock Channels 368 Alluvial Channels 368
- 13.6 Shaping Stream Valleys 370
 Base Level and Graded Streams 370
 Valley Deepening 371
 Valley Widening 372
 Incised Meanders and Stream Terraces 372
- 13.7 Depositional Landforms 374 Deltas 374 The Mississippi River Delta 375

Natural Levees **376** Alluvial Fans **376**

13.8 Floods and Flood Control 377 Types of Floods 377 Flood Control 378 Concepts in Review 380 Give It Some Thought 382

14 Groundwater 384

- 14.1 The Importance of Groundwater 386 Groundwater and the Hydrosphere 386 Geologic Importance of Groundwater 386 Groundwater: A Basic Resource 387
- 14.2 Groundwater and the Water Table 388 Distribution of Groundwater 388 The Water Table 388
- 14.3 Factors Influencing the Storage and Movement of Groundwater 391
 Porosity 391
 Permeability, Aquitards, and Aquifers 391
- How Groundwater Moves 392
 A Simple Groundwater Flow System 392
 Measuring Groundwater Movement 393
 Different Scales of Movement 393
- 14.5 Wells 394
- 14.6 Artesian Systems 395
- 14.7 Springs, Hot Springs, and Geysers 397
 Springs 397
 Hot Springs 397
 Geysers 398
- 14.8 Environmental Problems 400 Mining Groundwater 400



Subsidence **401** Saltwater Contamination **402** Groundwater Contamination **403**

14.9 The Geologic Work of Groundwater 404 Caverns 404 Karst Topography 405 Concepts in Review 408 Give It Some Thought 410

15 Glaciers and Glaciation 412

- **15.1 Glaciers: A Part of Two Basic Cycles 414** Valley (Alpine) Glaciers **415** Ice Sheets **415** Other Types of Glaciers **416**
- 15.2 Formation and Movement of Glacial Ice 417 Glacial Ice Formation 417 How Glaciers Move 417 Observing and Measuring Movement 418 Budget of a Glacier: Accumulation Versus Wastage 419
- 15.3 Glacial Erosion 421
 How Glaciers Erode 421
 Landforms Created by Glacial Erosion 422
- 15.4 Glacial Deposits 425 Glacial Drift 425 Moraines, Outwash Plains, and Kettles 426 Drumlins, Eskers, and Kames 428
- 15.5 Other Effects of Ice Age Glaciers 429 Crustal Subsidence and Rebound 429 Sea-Level Changes 429 Changes to Rivers and Valleys 430 Ice Dams Create Proglacial Lakes 431 Pluvial Lakes 431



15.6 The Ice Age 432
 Development of the Glacial Theory 432
 Causes of Ice Ages 433

 Concepts in Review 436
 Give It Some Thought 438

16 Deserts and Wind 440

- 16.1 Distribution and Causes of Dry Lands 442
 What Is Meant by Dry? 442
 Subtropical Deserts and Steppes 442
 Middle-Latitude Deserts and Steppes 443
- 16.2 Geologic Processes in Arid Climates 444
 Weathering 445
 The Role of Water 445
- 16.3 Basin and Range: The Evolution of a Desert Landscape 446

16.4 Wind Erosion 448 Transportation of Sediment by Wind 448 Erosional Features 450

16.5 Wind Deposits 452

 Sand Deposits 452
 Types of Sand Dunes 453
 Loess (Silt) Deposits 455
 Concepts in Review 456
 Give It Some Thought 457

17 Shorelines 458

- 17.1 The Shoreline: A Dynamic Interface 460 The Coastal Zone 460 Basic Features 460 Beaches 461
- 17.2 Ocean Waves 463 Wave Characteristics 463 Circular Orbital Motion 463 Waves in the Surf Zone 464
- 17.3 Shoreline Processes 465 Wave Erosion 465 Sand Movement on the Beach 465
- **17.4 Shoreline Features 468** Erosional Features **468** Depositional Features **468** The Evolving Shore **469**
- **17.5 Stabilizing the Shore 471** Hard Stabilization **472** Alternatives to Hard Stabilization **473**

Contents

- 17.6 Contrasting America's Coasts 474
 Atlantic and Gulf Coasts 474
 Pacific Coast 475
 Coastal Classification 476
- 17.7 Hurricanes: The Ultimate Coastal Hazard 477 Profile of a Hurricane 477 Hurricane Destruction 479

17.8 Tides 481 Causes of Tides 481 Monthly Tidal Cycle 482

Tidal Currents **482** Concepts in Review **483** Give It Some Thought **486**

18 Geologic Time 488

18.1 Creating a Time Scale: Relative Dating Principles 490 The Importance of a Time Scale 490 Numerical and Relative Dates 491 Principle of Superposition 491 Principle of Original Horizontality 491 Principle of Lateral Continuity 492 Principle of Cross-Cutting Relationships 492 Inclusions 492 Unconformities 493 Applying Relative Dating Principles 495

- 18.2 Fossils: Evidence of Past Life 496 Types of Fossils 496 Conditions Favoring Preservation 497
- 18.3 Correlation of Rock Layers 498 Correlation Within Limited Areas 498 Fossils and Correlation 500
- 18.4 Dating with Radioactivity 501
 Reviewing Basic Atomic Structure 501
 Radioactivity 501
 Half-Life 502
 Using Various Isotopes 503
 Dating with Carbon-14 504
- 18.5 The Geologic Time Scale 505
 Structure of the Time Scale 506
 Precambrian Time 506
 Terminology and the Geologic Time Scale 506
- 18.6 Determining Numerical Dates for Sedimentary Strata 507

Concepts in Review 508 Give It Some Thought 510

19 Earth's Evolution Through Geologic Time 512

19.1 Is Earth Unique? 514 The Right Planet 514 The Right Location 515 The Right Time 515 Viewing Earth's History 515 19.2 Birth of a Planet 517 From the Big Bang to Heavy Elements 517 From Planetesimals to Protoplanets 517 Earth's Early Evolution 517

19.3 Origin and Evolution of the Atmosphere and Oceans 519

Earth's Primitive Atmosphere **519** Oxygen in the Atmosphere **519** Evolution of the Oceans **520**

19.4 Precambrian History: The Formation of Earth's Continents 521 Earth's First Continents 522

The Making of North America **523** Supercontinents of the Precambrian **524**

- 19.5 Geologic History of the Phanerozoic: The Formation of Earth's Modern Continents 526 Paleozoic History 526 Mesozoic History 527 Cenozoic History 528
- 19.6 Earth's First Life 530 Origin of Life 530 Earth's First Life: Prokaryotes 532

19.7 Paleozoic Era: Life Explodes 532 Early Paleozoic Life-Forms 532 Vertebrates Move to Land 534



14 Contents

Reptiles: The First True Terrestrial Vertebrates **534** The Great Permian Extinction **534**

- 19.8 Mesozoic Era: Age of the Dinosaurs 536
 Gymnosperms: The Dominant Mesozoic Trees 536
 Reptiles: Dominating the Land, Sea, and Sky 536
 Demise of the Dinosaurs 537
- 19.9 Cenozoic Era: Age of Mammals 539
 From Reptiles to Mammals 539
 Marsupial and Placental Mammals 540
 Humans: Mammals with Large Brains and Bipedal

Locomotion **540** Large Mammals and Extinction **541 Concepts in Review 542**

Give It Some Thought 544

20 Global Climate Change 546

20.1 Climate and Geology 548 The Climate System 548 Climate–Geology Connections 548

20.2 Detecting Climate Change 549
 Seafloor Sediment—A Storehouse of Climate
 Data 550
 Oxygen Isotope Analysis 550
 Climate Change Recorded in Glacial Ice 551
 Tree Rings—Archives of Environmental History 551

20.3 Some Atmospheric Basics 553 Composition of the Atmosphere 553 Extent and Structure of the Atmosphere 554

Other Types of Proxy Data 552

20.4 Heating the Atmosphere 556

Energy from the Sun **556** The Paths of Incoming Solar Energy **557** Heating the Atmosphere: The Greenhouse Effect **557**

- 20.5 Natural Causes of Climate Change 558 Plate Movements and Orbital Variations 559 Volcanic Activity and Climate Change 559 Solar Variability and Climate 561
- 20.6 Human Impact on Global Climate 562 Rising CO₂ Levels 562 The Atmosphere's Response 563 The Role of Trace Gases 564
- 20.7 Climate-Feedback Mechanisms 566 Types of Feedback Mechanisms 566 Computer Models of Climate: Important yet Imperfect Tools 566
- 20.8 How Aerosols Influence Climate 567
- 20.9 Some Possible Consequences of Global Warming 568

Sea-Level Rise The Changing Arctic Increasing Ocean Acidity The Potential for "Surprises" Concepts in Review Give It Some Thought

APPENDIX

Metric and English Units Compared 576

GLOSSARY 577

INDEX 586

The 12th edition of *Essentials of Geology*, like its predecessors, is a college-level text for students taking their first and perhaps only course in geology. The text is intended to be a meaningful, non-technical survey for people with little background in science. Usually students are taking this class to meet a portion of their college's or university's general requirements.

In addition to being informative and up-to-date, a major goal of *Essentials of Geology* is to meet the need of beginning students for a readable and user-friendly text; a text that is a highly usable tool for learning the basic principles and concepts of geology.

New to the This Edition

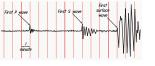
Preface

• New and expanded active learning path. *Essentials of Geology*, 12th edition, is designed for learning. Every chapter begins with *Focus on Concepts*. Each numbered learning objective corresponds to a major section in the chapter. The statements identify the knowledge and skills students should master by the end of the chapter, helping students prioritize key concepts.

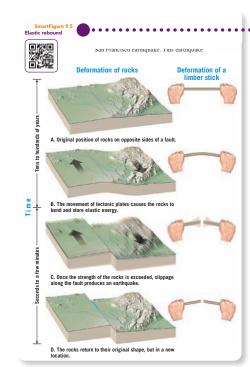


Give It Some Thought

- Do all earthquakes have a geological cause? What are the different ways in which human beings can contribute in earthquake generation?
- ⑦ The accompanying map shows the locations of many offset leagest earthquakes in the world since 1900. Refer to the map of Earth's plate boundaries in Figure 2.11 (page xx) and determine which type of plate boundary is most often associated with these destructive
- 3 Use the accompanying seismogram to answer the following
- questions:
 a. Which of the three types of seismic waves reached the seismo-graph first?
 b. What is the time interval between the arrival of the first P wave and the arrival of the first S wave?
- and the arrival of the first 5 wave? Use your answer from Question b and the travel-time graph in Figure 9.15 on page 247 to determine the distance from the sei mic station to the earthquake.
- Which of the three types of seismic waves had the highest amplitude when they reached the seismic station?



- (4) It is known that numbers on the Richter scale usually relate to the It is known that numbers on the Kichter scale usually relate to the amplitude of the largest earthquake waves. How can these number be converted to assess the actual amount of energy released due to the occurrence of such movements? What kind of measurements can be made directly from a seismogram?
- Which properties of primary waves (P Waves) and seco (S Waves) are used to determine the structure of Earth's (6) Actual destructiveness of an earthquake depends on two major fac-tors, other than energy release given by Richter magnitude. Name the other factors involved. Can any scale observe earth shaking effects?
- Concepts in Review. This new end-of-chapter feature is an important part of the book's revised active learning path. Each review is coordinated with the Focus on Concepts at the beginning of the chapter and with the numbered sections within the chapter. It is a readable and concise overview of key ideas, with photos, diagrams, and questions that also help students focus on important ideas and test their understanding of key concepts.



• Mobile Field Trips. Scattered through this new edition of Essentials of Geology are thirteen video field trips. On each trip, you will accompany geologist-pilot-photographer Michael Collier in the air and on the ground to see and learn about landscapes that relate to discussions in the chapter. These extraordinary field trips are accessed in the same way as SmartFigures: You simply scan a QR code that accompanies a figure in the chapter-usually one of Michael's outstanding photos.

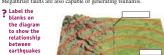
Each chapter concludes with *Give It Some Thought*. The questions and problems in this section challenge learners by involving them in activities that require higher-order thinking skills, such as application, analysis, and synthesis of material in the chapter. Some of the GIST problems are intended to develop an awareness of and appreciation for some of the Earth system's many interrelationships.

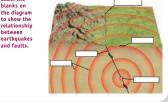
oncepts in Review Earthquakes and Earth's Interior

9.1 What Is an Earthquake?

- Sketch and describe the mechanism that generates most earthqu Key Terms: earthquake fault hypocenter (focus) enicenter seismic wave, elastic rebound, aftershock, foreshock, strike-slip fault, trans-form fault, fault creep, thrust fault, megathrust fault
- torm taut, ratu creep, must taut, megamust taut Earthquakes are caused by the sudden movement of blocks of rock on opposite sides of faults. The spot where the rock begins to slip is the hypocenter (or focus). Seismic waves radiate from this spot outward into the surrounding rock. The point on Earth's surface directly above the hypocenter is the epicenter. · Elastic rebound explains why most earthquakes happen: Rock is
- Elastic rebound explains why most earthquakes happen: Rock is deformed by movement of Earth's crust. However, frictional resistance keeps the fault locked in place, and the rock bends elastically. Strain builds up util it is greater than the resistance, and the blocks of rock suddenly slip, releasing the pent-up energy in the form of seismic wave. As elastic rebound occurs, the blocks of rock on either side of the fault return to their original shapes, but they are now in new positions. · Foreshocks are smaller earthquakes that precede larger earthquakes
- Aftershocks are smaller earthquakes that happen after large earthquake as the crust readjusts to the new, post-earthquake conditions.

- · Faults associated with plate boundaries are the source of most larg earthquakes
- · The San Andreas Fault in California is an example of a transform fault boundary capable of generating de-Subduction zones are marked by megathrust faults, large faults that are responsible for the largest earthquakes in recorded history Megathrust faults are also capable of generating tsunamis.





SmartFigures—art that teaches. SmartFigures. Essentials of Geology, 12th edition, has more than 100 of these figures distributed through each chapter. Just use your mobile device to scan the Quick Response (QR) code next to a SmartFigure, and the art comes alive. Each 2- to 3-minute feature, prepared and narrated by Professor Callan Bentley, is a mini-lesson that examines and explains the concepts illustrated by the figure. It is truly art that teaches.





- **Revised organization.** Earlier editions of this text had a more traditional chapter organization, in which the theory of plate tectonics was fully developed relatively late in the text. A major change to *Essentials of Geology*, 12th edition, is a reorganization in which this basic theory is presented in Chapter 2 to reflect the unifying role that plate tectonics plays in our understanding of planet Earth. With the basic framework of plate tectonics firmly established, we turn to discussions of Earth materials and the related processes of volcanism and metamorphism. This is followed by chapters that examine earthquakes, the origin and evolution of the ocean floor, and crustal deformation and mountain building. Along the way, students will clearly see the relationships among these phenomena and the theory of plate tectonics.
- An unparalleled visual program. In addition to more than 150 new, high-quality photos and satellite images, dozens of figures are new or have been redrawn by renowned geoscience illustrator Dennis Tasa. Maps and diagrams are frequently paired with photographs for greater effectiveness. Further, many new and revised figures have additional labels that narrate the process being illustrated and guide students as they examine the figures resulting in is a visual program that is clear and easy to understand.
- Significant updating and revision of content. A basic function of a college science textbook is to provide clear, understandable presentations that are accurate, engaging, and up-to-date. Our number-one goal is to keep *Essentials of Geology* current, relevant, and highly readable for beginning students. Every part of this text has been examined carefully with this goal in mind. Many discussions, case studies, and examples have been revised. The 12th edition represents perhaps the *most extensive and thorough revision* in the long history of this textbook.

Distinguishing Features Readability

The language of this text is straightforward and *written to be understood*. Clear, readable discussions with a minimum of technical language are the rule. The frequent headings and subheadings help students follow discussions and identify the important ideas presented in each chapter. In the 12th edition, we have continued to improve readability by examining chapter organization and flow and by writing in a more personal style. Significant portions of several chapters were substantially rewritten in an effort to make the material easier to understand.

Focus on Basic Principles and Instructor Flexibility

Although many topical issues are treated in the 12th edition of *Essentials of Geology*, it should be emphasized that the main focus of this new edition remains the same as the focus of each of its predecessors: to promote student understanding of basic principles. As much as possible, we have attempted to provide the reader with a sense of the observational techniques and reasoning processes that constitute the science of geology.

As in previous editions, we have designed most chapters to be self-contained so that material may be taught in a different sequence, according to the preference of the instructor or the dictates of the laboratory. Thus, an instructor who wishes to discuss erosional processes prior to earthquakes, plate tectonics, and mountain building may do so without difficulty.

A Strong Visual Component

Geology is highly visual, and art and photographs play a critical role in an introductory textbook. As in previous editions, Dennis Tasa, a gifted artist and respected geoscience illustrator, has worked closely with the authors to plan and produce the diagrams, maps, graphs, and sketches that are so basic to student understanding. The result is art that is clearer and easier to understand than ever before.

Our aim is to get *maximum effectiveness* from the visual component of the text. Michael Collier, an award-winning geologist–pilot–photographer, aided greatly in this quest. As you read through this text, you will see dozens of his extraordinary aerial photographs. His contributions truly help bring geology alive for the reader.

17

The Teaching and Learning Package For the Instructor

Pearson continues to improve the instructor resources for this text, with the goal of providing dynamic teaching aids and saving you time in preparing for your classes.

Instructor's Resource Centre

The IRC provides an integrated collection of resources designed to help instructors make efficient and effective use of their time. It features:

- **PowerPointTM Presentations:** The presentation contains a complete and customizable lecture outline with supporting art.
- The Geoscience Animation Library including more than 100 animations that illustrate many difficult-to-visualize topics of geology. Created through a unique collaboration among five of Pearson's leading geoscience authors, these animations represent a significant step forward in lecture presentation aids.
- **Instructor's Manual** containing learning objectives, chapter outlines, answers to end-of-chapter questions, and suggested short demonstrations to spice up your lecture. The Test Bank incorporates art and averages 75 multiple-choice, true/false, short-answer, and critical thinking questions per chapter.
- **TestGen:** An electronic version of the Testbank that allows you to customize and manage your tests. Testbank is also available in Microsoft Word.

Course Management

Pearson offers instructor and student media for this 12th edition of *Essentials of Geology* in formats compatible with Blackboard and other course management platforms. Contact your local Pearson representative for more information.

For the Student

The student resources to accompany *Essentials of Geology*, 12th edition, have been further refined, with the goal of focusing the students' efforts and improving their understanding of the concepts of geology.

- SmartFigures
- Mobile Field Trips
- Web Links
- Glossary

Acknowledgments

Writing a college textbook requires the talents and cooperation of many people. It is truly a team effort, and the authors are fortunate to be part of an extraordinary team at Pearson Education. In addition to being great people to work with, all are committed to producing the best textbooks possible. Special thanks to our geology editor, Andy Dunaway, who invested a great deal of time, energy, and effort in this project. We appreciate his enthusiasm, hard work, and quest for excellence. We also appreciate our conscientious project manager, Crissy Dudonis, whose job it was to keep track of all that was going on-and a lot was going on. The text's new design resulted from the creative talents of Derek Bacchus and his team. We think it is a job well done. As always, our marketing manager, Maureen McLaughlin, provided helpful advice and many good ideas. Essentials of Geology, 12th edition, was truly improved with the help of our developmental editor Jonathan Cheney. Many thanks. It was the job of the production team, led by Heidi Allgair at Cenveo® Publisher Services, to turn our manuscript into a finished product. The team also included copyeditor Kitty Wilson, compositor Annamarie Boley, proofreader Heather Mann, and photo researcher Kristin Piljay. We think these talented people did great work. All are true professionals, with whom we are very fortunate to be associated.

The authors owe special thanks to three people who were very important contributors to this project:

- Working with Dennis Tasa, who is responsible for all of the text's outstanding illustrations, is always special for us. He has been part of our team for more than 30 years. We not only value his artistic talents, hard work, patience, and imagination but his friendship as well.
- As you read this text, you will see dozens of extraordinary photographs by Michael Collier. Most are aerial shots taken from his nearly 60-year-old Cessna 180. Michael was also responsible for preparing the remarkable Mobile Field Trips that are scattered through the text. Among his many awards is the American Geological Institute Award for Outstanding contribution to the Public Understanding of Geosciences. We think that Michael's photographs and field trips are the next best thing to being there. We were very fortunate to have had Michael's assistance on *Essentials of Geology*, 12th edition. Thanks, Michael.
- Callan Bentley has been an important addition to the *Essentials* of *Geology* team. Callan is an assistant professor of geology at Northern Virginia Community College in Annandale, where he has been honored many times as an outstanding teacher. He is a frequent contributor to *Earth* magazine and is author of the popular geology blog *Mountain Beltway*. Callan was responsible

for preparing the SmartFigures that appear throughout the text. As you take advantage of these outstanding learning aids, you will hear his voice explaining the ideas. Callan also contributed to the Concepts in Review feature found at the end of each chapter. We appreciate Callan's contributions to this new edition of *Essentials of Geology*.

Great thanks also go to those colleagues who prepared in-depth reviews. Their critical comments and thoughtful input helped guide our work and clearly strengthened the text. Special thanks to:

Tania Anders, Texas A&M University–Corpus Christi Jamie Barnes, University of Texas–Austin David Bradley, Georgia Southern University Alan Coulson, Clemson University Sarah de la Rue, Purdue–Calumet Noah Fay, Pima Community College Thomas Gerber, Indiana University of PA Wayne Henderson, California State University– Fullerton Edgar Kessler, Northampton Community College

Katherine Knierim, University of Arkansas Sam Matson, Boise State University Charles Merguerian, Hofstra University Stephen Moysey, Clemson University Jodi Ryder, Central Michigan University Robert Shuster, University of Nebraska–Omaha Gordana Vlahovic, North Carolina Central College Merry Wilson, Scottsdale Community College Chris Woltemade, Shippensburg University Adam Woods, California State University–Fullerton Sally Zellers, University of Central Missouri James Zollweg, Boise State University

Last, but certainly not least, we gratefully acknowledge the support and encouragement of our wives, Nancy Lutgens and Joanne Bannon. Preparation of *Essentials of Geology*, 12th edition, would have been far more difficult without their patience and understanding.

Fred Lutgens Ed Tarbuck

Pearson would like to thank and acknowledge Ananya Biswas (Bengal Engineering and Science University) for her contributions to the Global Edition, and Christopher Satow (Kingston University), Arun Singh (IIT Kharagpur), Read Brown Mapeo (University of Botswana) and Supriyo Kumar Das (Presidency University) for their guidance and recommendations to improve the global content.

An Introduction to Geology

ocus on Concepts

Each statement represents the primary learning objective for the corresponding major heading within the chapter. After you complete the chapter, you should be able to:

- **1.1** Distinguish between physical and historical geology and describe the connections between people and geology.
- **1.2** Summarize early and modern views on how change occurs on Earth and relate them to the prevailing ideas about the age of Earth.
- **1.3** Discuss the nature of scientific inquiry, including the construction of hypotheses and the development of theories.
- **1.4** List and describe Earth's four major spheres.
- **1.5** Define *system* and explain why Earth is considered to be a system.
- **1.6** Outline the stages in the formation of our solar system.
- **1.7** Describe Earth's internal structure.
- **1.8** Sketch, label, and explain the rock cycle.
- **1.9** List and describe the major features of the continents and ocean basins.

The view from Toroweap Overlook along the North Rim of Arizona's Grand Canyon National Park. (Photo by Michael Collier) **The spectacular eruption of a volcano,** the terror brought by an earthquake, the magnificent scenery of a mountain range, and the destruction created by a landslide or flood are all subjects for a geologist. The study of geology deals with many fascinating and practical questions about our physical environment. What forces produce mountains? Will there soon be a major earthquake in California? What was the Ice Age like, and will there be another? How were ore deposits formed? Where should we search for water? Will plentiful oil be found if a well is drilled in a particular location? Geologists seek to answer these and many other questions about Earth, its history, and its resources.

Figure 1.1 Internal and external processes The processes that operate beneath and upon Earth's surface are an important focus of physical geology. (Volcano photo by Lucas Jackson/Reuters; glacier photo by Michael Collier)



Geology: The Science of Earth

Distinguish between physical and historical geology and describe the connections between people and geology.



Internal processes are those that occur beneath Earth's surface. Sometimes they lead to the formation of major features at the surface.



mountains in Alaska.

The subject of this text is **geology**, from the Greek geo (Earth) and logos (discourse). Geology is the science that pursues an understanding of planet Earth. Understanding Earth is challenging because our planet is a dynamic body with many interacting parts and a complex history. Throughout its long existence, Earth has been changing. In fact, it is changing as you read this page and will continue to do so into the foreseeable future. Sometimes the changes are rapid and violent, as when landslides or volcanic eruptions occur. Just as often, change takes place so slowly that it goes unnoticed during a lifetime. Scales of size and space also vary greatly among the phenomena that geologists study. Sometimes geologists must focus on phenomena that are microscopic, and at other times they must deal with features that are continental or global in scale.

Physical and Historical Geology

Geology is traditionally divided into two broad areasphysical and historical. Physical geology, which is the primary focus of this book, examines the materials composing Earth and seeks to understand the many processes that operate beneath and upon its surface (Figure 1.1). The aim of historical geology, on the other hand, is to understand the origin of Earth and its development through time. Thus, it strives to establish an orderly chronological arrangement of the multitude of physical and biological changes that have occurred in the geologic past. The study of physical geology logically precedes the study of Earth history because we must first understand how Earth works before we attempt to unravel its past. It should also be pointed out that physical and historical geology are divided into many areas of specialization. Every chapter of this book represents one or more areas of specialization in geology.



Figure 1.2 In the field and in the lab Geology not only involves outdoor fieldwork but work in the laboratory as well. (Photo by British Antarctic Survey/Science Source)

tsunamis, earthquakes, and landslides. Of course, geologic hazards are *natural* processes. They become hazards only when people try to live where these processes occur.

According to the United Nations, in 2008, for the first time, more people lived in cities than in rural areas. This global trend toward urbanization concentrates millions of people into megacities, many of which are vulnerable to natural hazards. Coastal sites are becoming more vulnerable because development often destroys natural defenses such as wetlands and sand dunes. In addition, there is a growing threat associated with human influences on the Earth system; one example is sea-level rise that is linked to global climate change. Some megacities are exposed to seismic (earthquake) and volcanic

hazards where inappropriate land use and poor construction practices, coupled with rapid population growth, are increasing vulnerability.

Resources are another important focus of geology that is of great practical value to people. They include water and soil, a great variety of metallic and nonmetallic

Did You Know?

Each year an average American requires huge quantities of Earth materials. Imagine receiving your annual share in a single delivery. A large truck would pull up to your home and unload 12,965 lb of stone, 8945 lb of sand and gravel, 895 lb of cement, 395 lb of salt, 361 lb of phosphate, and 974 lb of other nonmetals. In addition, there would be 709 lb of metals, including iron, aluminum, and copper.

Did You Know?

It took until about the year 1800 for the world population to reach 1 billion. By 1927, the number had doubled to 2 billion. According to United Nations estimates, world population reached 7 billion in late October 2011. We are currently adding about 80 million people per year to the planet.

outdoors—and rightly so. A great deal of geology is based on observations, measurements, and experiments conducted in the field. But geology is also done in the laboratory, where, for example, the analysis of minerals and rocks provides insights into many basic processes and the microscopic study of fossils unlocks

Geology is perceived as a science that is done

clues to past environments (**Figure 1.2**). Frequently, geology requires an understanding and application of knowledge and principles from physics, chemistry, and biology. Geology is a science that seeks to expand our knowledge of the natural world and our place in it.

Geology, People, and the Environment

The primary focus of this book is to develop an understanding of basic geologic principles, but along the way we will explore numerous important relationships between people and the natural environment. Many of the problems and issues addressed by geology are of practical value to people.

Natural hazards are a part of living on Earth. Every day they adversely affect millions of people worldwide and are responsible for staggering damages (**Figure 1.3**). Among the hazardous Earth processes that geologists study are volcanoes, floods, Figure 1.3 Earthquake destruction Geologic hazards are natural processes. They become hazards only

when people try to live where these processes occur. (Photo by Yasuyoshi Chiba/AFP/Getty Images/Newscom)





Figure 1.4 Drilling for

oil Energy and mineral resources represent an important link between people and geology. Petroleum provides more than 36 percent of U.S. energy consumption. (Photo by Peter Bowater/Science Source)

minerals, and energy (**Figure 1.4**). Together they form the very foundation of modern civilization. Geology deals not only with the formation and occurrence of these vital resources but also with maintaining supplies and with the environmental impact of their extraction and use.

Geologic processes clearly have an impact on people. In addition, we humans can dramatically influence geologic processes. For example, river flooding is natural, but the magnitude and frequency of flooding can be affected significantly by human activities such as clearing forests, building cities, and constructing dams. Unfortunately, natural systems do not always adjust to artificial changes in ways that we can anticipate. Thus, an alteration to the environment that was intended to benefit society sometimes has the opposite effect.

At appropriate places throughout this book, you will have opportunities to examine different aspects of our relationship with the physical environment. It will be rare to find a chapter that does not address some aspect of natural hazards, environmental issues, or resources. Significant parts of some chapters provide the basic geologic knowledge and principles needed to understand environmental problems.

Concept Checks 1.1

- Name and distinguish between the two broad subdivisions of geology.
- 2 List at least three different geologic hazards.
- 3 Aside from geologic hazards, describe another important connection between people and geology.

1.2 The Development of Geology

Summarize early and modern views on how change occurs on Earth and relate them to the prevailing ideas about the age of Earth.

The nature of our Earth—its materials and processes—has been a focus of study for centuries. Writings about such topics as fossils, gems, earthquakes, and volcanoes date back to the early Greeks, more than 2300 years ago.

Certainly the most influential Greek philosopher was Aristotle. Unfortunately, Aristotle's explanations about the natural world were not based on keen observations and experiments. Instead, they were arbitrary pronouncements. He believed that rocks were created under the "influence" of the stars and that earthquakes occurred when air crowded into the ground, was heated by central fires, and escaped explosively. When confronted with a fossil fish, he explained that "a great many fishes live in the earth motionless and are found when excavations are made." Although Aristotle's explanations may have been adequate for his day, they unfortunately continued to be viewed as authoritative for many centuries, thus inhibiting the acceptance of more up-to-date ideas. After the Renaissance of the 1500s, however, more people became interested in finding answers to questions about Earth.

Catastrophism

In the mid-1600s, James Ussher, Anglican Archbishop of Armagh, Primate of all Ireland, published a major work that had immediate and profound influences. A respected scholar of the Bible, Ussher constructed a chronology of human and Earth history in which he calculated that Earth was only a few thousand years old, having been created in 4004 B.C. Ussher's treatise earned widespread acceptance among Europe's scientific and religious leaders, and his chronology was soon printed in the margins of the Bible itself.

During the seventeenth and eighteenth centuries, Western thought about Earth's features and processes was strongly influenced by Ussher's calculation. The result was a guiding doctrine called **catastrophism**. Catastrophists believed that Earth's landscapes were shaped primarily by great catastrophes. Features such as mountains and canyons, which today we know take great spans of time to form, were explained as having been produced by sudden and often worldwide disasters produced by unknowable causes that no longer operate. This philosophy was an attempt to fit the rates of Earth processes to the then-current ideas on the age of Earth.

The Birth of Modern Geology

Against the backdrop of Aristotle's views and an Earth created in 4004 B.C., a Scottish physician and gentleman farmer named James Hutton published *Theory of the Earth* in 1795. In this work, Hutton put forth a fundamental principle that is a pillar of geology today: **uniformitarianism**. It states that the *physical, chemical, and biological laws that operate today*

have also operated in the geologic past. This means that the forces and processes that we observe presently shaping our planet have been at work for a very long time. Thus, to understand ancient rocks, we must first understand presentday processes and their results. This idea is commonly stated as *the present is the key to the past*.

Prior to Hutton's *Theory of the Earth*, no one had effectively demonstrated that geologic processes occur over extremely long periods of time. However, Hutton persuasively argued that forces that appear small can, over long spans of time, produce effects that are just as great as those resulting from sudden catastrophic events. Unlike his predecessors, Hutton carefully cited verifiable observations to support his ideas.

For example, when Hutton argued that mountains are sculpted and ultimately destroyed by weathering and the work of running water and that their wastes are carried to the oceans by processes that can be observed, he said, "We have a chain of facts which clearly demonstrate . . . that the materials of the wasted mountains have traveled through the rivers"; and further, "There is not one step in all this progress . . . that is not to be actually perceived." He then went on to summarize this thought by asking a question and immediately providing the answer: "What more can we require? Nothing but time."

Mobile Field Trip

Grand Canyon rocks span more than 1.5 billion years of Earth history.

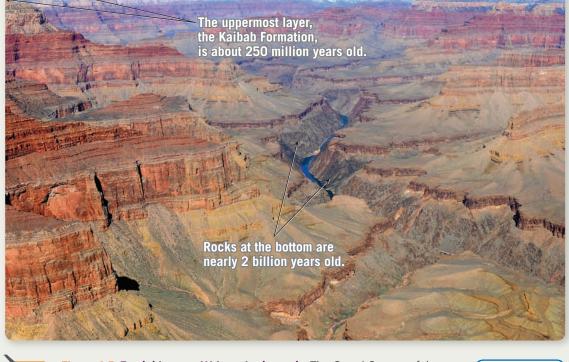


Figure 1.5 Earth history—Written in the rocks The Grand Canyon of the Colorado River in northern Arizona. (*Photo by Dennis Tasa*)



Geology Today

Today the basic tenets of uniformitarianism are just as viable as in Hutton's day. Indeed, today we realize more strongly than ever before that the present gives us insight into the past and that the physical, chemical, and biological laws that govern geologic processes remain unchanging through time. However, we also understand that the doctrine should not be taken too literally. To say that geologic processes in the past were the same as those occurring today is not to suggest that they have always had the same relative importance or that they have operated at precisely the same rate. Moreover, some important geologic processes are not currently observable, but evidence that they occur is well established. For example, we know that Earth has experienced impacts from large meteorites even though we have no human witness accounts of those impacts. Nevertheless, such events have altered Earth's crust, modified its climate, and strongly influenced life on the planet.

The acceptance of uniformitarianism meant the acceptance of a very long history for Earth. Although Earth processes vary in intensity, they still take a very long time to create or destroy major landscape features. The Grand Canyon provides a good example (Figure 1.5).

Did You Know?

Shortly after Archbishop Ussher determined an age for Earth, another biblical scholar, Dr. John Lightfoot of Cambridge, felt he could be even more specific. He wrote that Earth was created "on the 26th of October 4004 BC at 9 o'clock in the morning." (As quoted in William L. Stokes, *Essentials of Earth History*, Prentice Hall, Inc. 1973, p. 20.) An Introduction to Geology

Did You Know?

Estimates indicate that erosional processes are lowering the North American continent at a rate of about 3 cm per 1000 years. At this rate, it would take 100 million years to level a 3000 m (10,000 ft) high peak. The rock record contains evidence which shows that Earth has experienced many cycles of mountain building and erosion. Concerning the ever-changing nature of Earth through great expanses of geologic time, Hutton made a statement that was to become his most famous. In concluding his classic 1788 paper published in the *Transactions of the Royal Society of Edinburgh*, he stated, "The results, therefore, of our present enquiry is, that we find no vestige of a beginning—no prospect of an end."

In the chapters that follow, we will be examining the materials that compose our planet and the processes that modify it. It is important to remember that, although many features of our physical landscape may seem to be unchanging over the decades we observe them, they are nevertheless changing—but on time scales of hundreds, thousands, or even many millions of years.

The Magnitude of Geologic Time

Among geology's important contributions to human knowledge is the discovery that Earth has a very long and complex history. Although Hutton and others recognized that geologic time is exceedingly long, they had no methods to accurately determine the age of Earth. Early time scales simply placed the events of Earth history in the proper sequence or order, without knowledge of how long ago in years they occurred.

Today our understanding of radioactivity allows us to accurately determine numerical dates for rocks that represent important events in Earth's distant past (**Figure 1.6**). For example, we know that the dinosaurs died out about 65 million years ago. Today the age of Earth is put at about 4.6 billion years. Chapter 18 is devoted to a much more complete discussion of geologic time and the geologic time scale.

The concept of geologic time is new to many nongeologists. People are accustomed to dealing with increments of time that are measured in hours, days, weeks, and years. Our history books often examine events over spans of centuries, but even a century is difficult to appreciate fully. For most of us, someone or something that is 90 years old is *very old*, and a 1000-year-old artifact is *ancient*.

By contrast, those who study geology must routinely deal with vast time periods—millions or billions (thousands of millions) of years. When viewed in the context of Earth's 4.6-billion-year history, a geologic event that occurred 100 million years ago may be characterized as "recent" by a geologist, and a rock sample that has been dated at 10 million years may be called "young." An appreciation for the magnitude of geologic time is important in the study of geology because many processes are so gradual that vast spans of time are needed

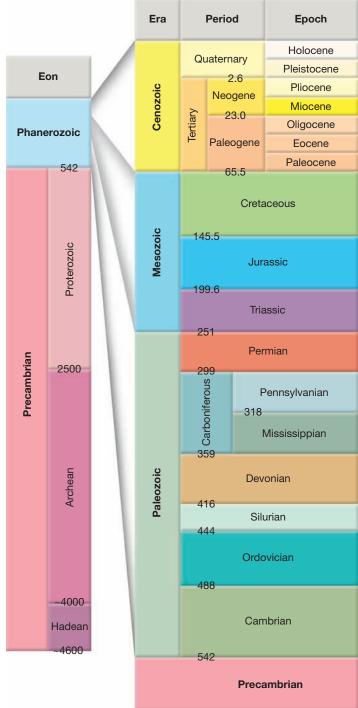
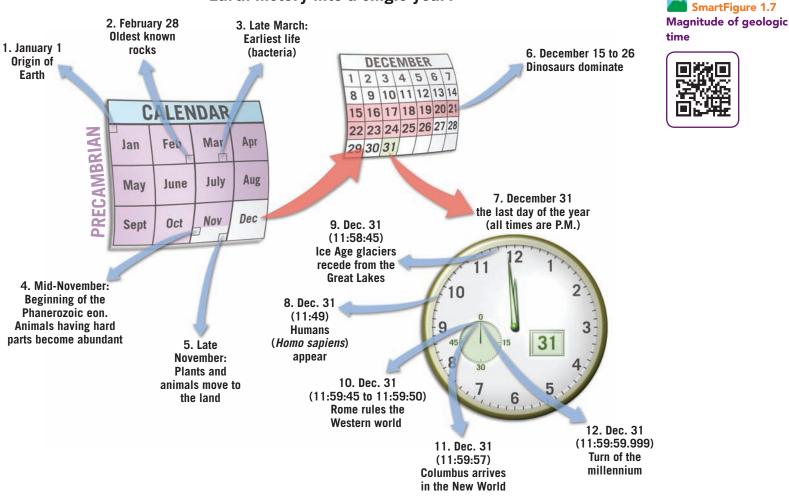


Figure 1.6 Geologic time scale: A basic reference The time scale divides the vast 4.6-billion-year history of Earth into eons, eras, periods, and epochs. Numbers on the time scale represent time in millions of years before the present. The Precambrian accounts for more than 88 percent of geologic time.

What if we compress the 4.6 billion years of Earth history into a single year?



before significant changes occur. How long is 4.6 billion years? If you were to begin counting at the rate of one number per second and continued 24 hours a day, 7 days a week and never stopped, it would take about two lifetimes (150 years) to reach 4.6 billion! **Figure 1.7** provides another interesting way of viewing the expanse of geologic time.

The foregoing is just one of many analogies that have been conceived in an attempt to convey the magnitude of geologic time. Although helpful, all of them, no matter how clever, only begin to help us comprehend the vast expanse of Earth history.

Concept Checks 1.2

- 1) Describe Aristotle's influence on geology.
- (2) Contrast catastrophism and uniformitarianism. How did each view the age of Earth?
- (3) How old is Earth?
- (4) Refer to Figure 1.6 and list the eon, era, period, and epoch in which we live.
- (5) Why is an understanding of the magnitude of geologic time important for a geologist?

1.3 The Nature of Scientific Inquiry

Discuss the nature of scientific inquiry, including the construction of hypotheses and the development of theories.

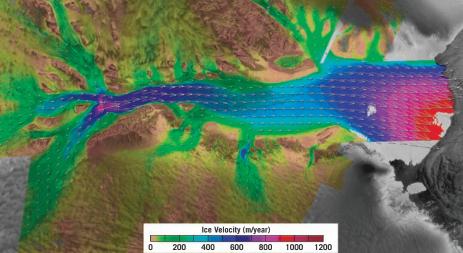
As members of a modern society, we are constantly reminded of the benefits derived from science. But what exactly is the nature of scientific inquiry? Science is a process of producing knowledge. The process depends both on making careful observations and on creating explanations that make sense of the observations. Developing an 27

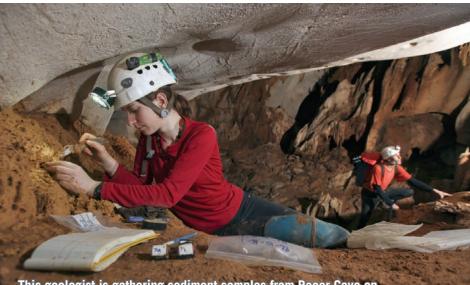
understanding of how science is done and how scientists work is an important theme that appears throughout this book. You will explore the difficulties in gathering data and some of the ingenious methods that have been developed to overcome these difficulties. You will also see many examples of how hypotheses are formulated and tested, as well as learn about the evolution and development of some major scientific theories.

All science is based on the assumption that the natu-

ral world behaves in a consistent and predictable manner that is comprehensible through careful, systematic study. Figure 1.8 Observation The overall goal of science is to discover the underlying and measurement Scienpatterns in nature and then to use that knowledge to make tific facts are gathered in predictions about what should or should not be expected, many ways. (Satellite image given certain facts or circumstances. For example, by by NASA; photo by Robbie

Instruments aboard satellites provide detailed information about the movement of Antarctica's Lambert Glacier. Such data are basic to understanding glacier behavior.





This geologist is gathering sediment samples from Racer Cave on the island of Borneo.

knowing how oil deposits form, geologists are able to predict the most favorable sites for exploration and, perhaps as importantly, how to avoid regions that have little or no potential.

The development of new scientific knowledge involves some basic logical processes that are universally accepted. To determine what is occurring in the natural world, scientists collect scientific facts through observation and measurement (Figure 1.8). The facts that are collected often seek to answer well-defined questions about the natural world. Because some error is inevitable, the accuracy of a particular measurement or observation is always open to question. Nevertheless, these data are essential to science and serve as a springboard for the development of scientific theories.

Hypothesis

Once facts have been gathered and principles have been formulated to describe a natural phenomenon, investigators try to explain how or why things happen in the manner observed. They often do this by constructing a tentative (or untested) explanation, which is called a scientific hypothesis. It is best if an investigator can formulate more than one hypothesis to explain a given set of observations. If an individual scientist is unable to devise multiple hypotheses, others in the scientific community will almost always develop alternative explanations. A spirited debate frequently ensues. As a result, extensive research is conducted by proponents of opposing hypotheses, and the results are made available to the wider scientific community in scientific journals.

Before a hypothesis can become an accepted part of scientific knowledge, it must pass objective testing and analysis. If a hypothesis cannot be tested, it is not scientifically useful, no matter how interesting it might seem. The verification process requires that *predictions* be made, based on the hypothesis being considered, and that the predictions be tested through comparison against objective observations of nature. Put another way, hypotheses must fit observations other than those used to formulate them in the first place. Hypotheses that fail rigorous testing are ultimately discarded. The history of science is littered with discarded hypotheses. One of the best known is the Earthcentered model of the universe-a proposal that was supported by the apparent daily motion of the Sun, Moon, and stars around Earth. As the mathematician Jacob Bronowski so ably stated, "Science is a great many things, but in the end they all return to this: Science is the acceptance of what works and the rejection of what does not."

Theory

When a hypothesis has survived extensive scrutiny and when competing hypotheses have been eliminated, a hypothesis may be elevated to the status of a scientific theory. In everyday language, we may say, "That's

Shone/Science Source)

only a theory." But a scientific theory is a well-tested and widely accepted view that the scientific community agrees best explains certain observable facts. Some theories that are extensively documented and extremely well supported are comprehensive in scope. For example, the theory of plate tectonics provides a framework for understanding the origin of mountains, earthquakes, and volcanic activity. In addition, plate tectonics explains the evolution of the continents and the ocean basins through time—ideas that are explored in some detail in Chapters 2, 10, and 11.

Scientific Methods

The process just described, in which researchers gather facts through observations and formulate scientific hypotheses and theories, is called the *scientific method*. Contrary to popular belief, the scientific method is not a standard recipe that scientists apply in a routine manner to unravel the secrets of our natural world. Rather, it is an endeavor that involves creativity and insight. Rutherford and Ahlgren put it this way: "Inventing hypotheses or theories to imagine how the world works and then figuring out how they can be put to the test of reality is as creative as writing poetry, composing music, or designing skyscrapers."*

There is not a fixed path that scientists always follow that leads unerringly to scientific knowledge. However, many scientific investigations involve the following:

- A question is raised about the natural world.
- Scientific data that relate to the question are collected.
- Questions that relate to the data are posed, and one or more working hypotheses are developed that may answer these questions.
- Observations and experiments are developed to test the hypotheses.
- The hypotheses are accepted, modified, or rejected, based on extensive testing.
- Data and results are shared with the scientific community for critical examination and further testing.

Some scientific discoveries result from purely theoretical ideas that stand up to extensive examination. Some researchers use high-speed computers to create models that simulate what is happening in the "real" world. These models are useful when dealing with natural processes that occur on very long time scales or that take place in extreme or inaccessible locations. Still other scientific advancements are made when a totally unexpected happening occurs during an experiment. These serendipitous discoveries are more than pure luck, for as the nineteenth-century French scientist Louis Pasteur said, "In the field of observation, chance favors only the prepared mind."

Scientific knowledge is acquired through several avenues, so it might be best to describe the nature of scientific inquiry as the methods of science rather than as the scientific method. In addition, it should always be remembered that even the most compelling scientific theories are still simplified explanations of the natural world.

Plate Tectonics and Scientific Inquiry

There are many opportunities in the pages of this book to develop and reinforce your understanding of how science works and, in particular, how the science of geology works. You will learn about the methods involved in gathering data and develop a sense of the observational techniques and reasoning processes used by geologists. Chapter 2 is an excellent example.

During the past several decades, a great deal has been learned about the workings of our dynamic planet. This period has seen an unequaled revolution in our understanding of Earth. The revolution began in the early part of the twentieth century, with the radical proposal of *continental drift*—the idea that the continents move about the face of the planet. This hypothesis contradicted the established view that the continents and ocean basins are permanent and stationary features on the face of Earth. For that reason, the notion of drifting continents was received with great skepticism and even ridicule. More than 50 years passed before enough data were gathered to transform this controversial hypothesis into a sound theory that wove together the basic processes known to operate on Earth. The theory that finally emerged, called the *theory of plate tectonics*, provided geologists with the first comprehensive model of Earth's internal workings.

As you read Chapter 2, you will not only gain insights into the workings of our planet, you will also see an excellent example of the way geologic "truths" are uncovered and reworked.

Concept Checks 1.3

- How is a scientific hypothesis different from a scientific theory?
- (2) Summarize the basic steps followed in many scientific investigations.
- 3 What explains the fact that continental drift is considered a hypothesis but plate tectonics is considered a theory?

Did You Know?

A scientific *law* is a basic principle that describes a particular behavior of nature that is generally narrow in scope and can be stated briefly—often as a simple mathematical equation.

Did You Know?

In 1492 when Columbus set sail, many Europeans thought that Earth was flat and that Columbus would sail off the edge. However, more than 2000 years earlier, ancient Greeks realized that Earth was spherical because it always cast a curved shadow on the Moon during a lunar eclipse. In fact, Eratosthenes (276-194 B.c.) calculated Earth's circumference and obtained a value close to the modern measurement of 40,075 km (24,902 mi).

1.4 Earth's Spheres

List and describe Earth's four major spheres.

The images in **Figure 1.9** are considered to be classics because they let humanity see Earth differently than ever before. These early views profoundly altered our conceptualizations of Earth and remain powerful images decades after they were first viewed. Seen from space, Earth is breathtaking in its beauty and startling in its solitude. The photos remind us that our home is, after

Figure 1.9 Two classic views of Earth from space (NASA)





This image taken from *Apollo* 17 in December 1972 is perhaps the first to be called "The Blue Marble." The dark blue ocean and swirling cloud patterns remind us of the importance of the oceans and atmosphere. all, a planet—small, selfcontained, and in some ways even fragile.

As we look closely at our planet from space, it becomes apparent that Earth is much more than rock and soil. In fact, perhaps the most conspicuous features in Figure 1.9 are swirling clouds suspended above the surface of the vast global ocean. These features emphasize the importance of water on our planet.

The closer view of Earth from space shown in Figure 1.9 helps us appreciate why the physical environment is traditionally divided into three major parts: the water portion of our planet, the *hydrosphere*; Earth's gaseous envelope, the *atmosphere*; and, of course, the solid Earth, or *geo-sphere*. It needs to be emphasized that our environment is highly integrated and not dominated by rock, water, or air alone. Rather, it is characterized by continuous interactions as air comes in contact with rock, rock with water, and water with air. Moreover, the *biosphere*, which is the totality of all plant and animal life on our planet, interacts with each of the three physical realms and is an equally integral part of the planet. Thus, Earth can be thought of as consisting of four major spheres: the hydrosphere, atmosphere, geosphere, and biosphere.

The interactions among Earth's four spheres are incalculable. **Figure 1.10** provides us with one easy-to-visualize example. The shoreline is an obvious meeting place for rock, water, and air. In this scene, ocean waves that were created by the drag of air moving across the water are breaking against the rocky shore. The force of the water can be powerful, and the erosional work that is accomplished can be great.

Hydrosphere

Earth is sometimes called the *blue* planet. Water, more than anything else, makes Earth unique. The **hydrosphere** is a dynamic mass of water that is continually on the move, evaporating from the oceans to the atmosphere, precipitat-

Figure 1.10 Interactions among Earth's spheres The shoreline is one obvious interface—a common boundary where different parts of a system interact. In this scene, ocean waves (hydrosphere) that were created by the force of moving air (atmosphere) break against a rocky shore (geosphere). The force of the water can be powerful, and the erosional work that is accomplished can be great. (*Photo by Michael Collier*)



ing to the land, and running back to the ocean again. The global ocean is certainly the most prominent feature of the hydrosphere, blanketing nearly 71 percent of Earth's surface to an average depth of about 3800 meters (12,500 feet). It accounts for about 97 percent of Earth's water (Figure 1.11). However, the hydrosphere also includes the freshwater found underground and in streams, lakes, and glaciers. Moreover, water is an important component of all living things.

Although these latter sources constitute just a tiny fraction of the total, they are much more important than their meager percentage indicates. In addition to providing the freshwater that is so vital to life on land, streams, glaciers, and groundwater are responsible for sculpting and creating many of our planet's varied landforms.

Atmosphere

Earth is surrounded by a life-giving gaseous envelope called the **atmosphere** (Figure 1.12). When we watch a high-flying jet plane cross the sky, it seems that the atmosphere extends upward for a great distance. However, when compared to the thickness (radius) of the solid Earth (about 6400 kilometers [4000 miles]), the atmosphere is a very shallow layer. Despite its modest

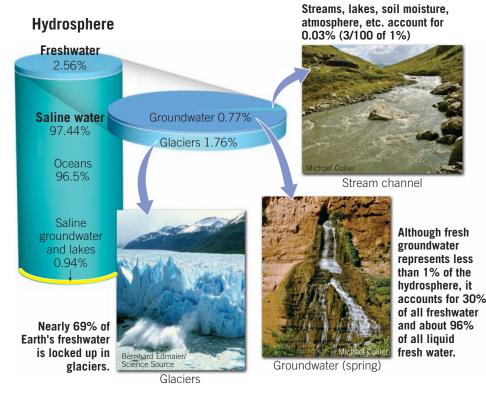
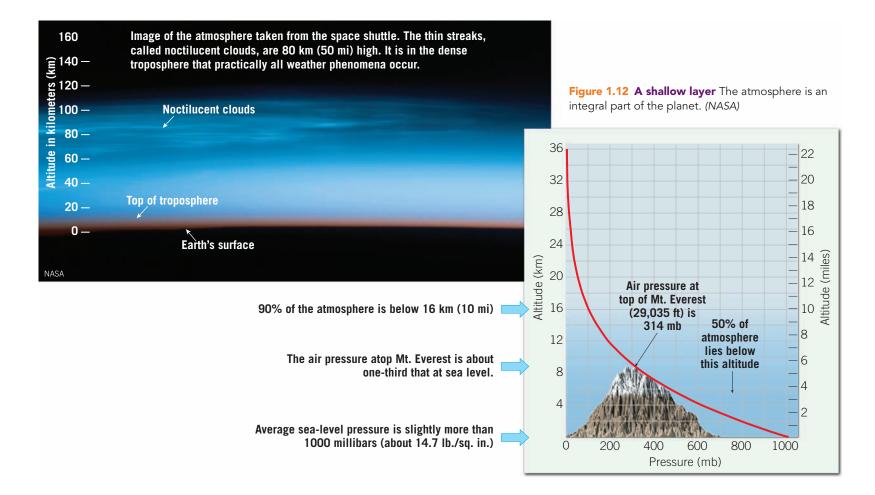


Figure 1.11 The water planet Distribution of water in the hydrosphere.



An Introduction to Geology

Did You Know?

The volume of ocean water is so large that if Earth's solid mass were perfectly smooth (level) and spherical, the oceans would cover Earth's entire surface to a uniform depth of more than 2000 m (1.2 mi).

dimensions, this thin blanket of air is an integral part of the planet. It not only provides the air that we breathe but also protects us from the Sun's intense heat and dangerous ultraviolet radiation. The energy exchanges that continually occur between the atmosphere and Earth's surface and between the atmosphere and space produce the effects we call weather and climate. Climate has a strong influence on the nature and intensity of Earth's external processes. When climate changes, these processes respond.

If, like the Moon, Earth had no atmosphere, our planet would be lifeless, and many of the processes and interactions that make the surface such a dynamic place could not operate. Without weathering and erosion, the face of our planet might more closely resemble the lunar surface, which has not changed appreciably in nearly 3 billion years.

Biosphere

The **biosphere** includes all life on Earth (Figure 1.13). Ocean life is concentrated in the sunlit surface waters of the sea. Most life on land is also concentrated near the surface, with tree roots and burrowing animals reaching a few meters underground and flying insects and birds reaching a kilometer or so into the atmosphere. A sur-

prising variety of life-forms are also adapted to extreme environments. For example, on the ocean floor, where pressures are extreme and no light penetrates, there are places where vents spew hot, mineral-rich fluids that support communities of exotic life-forms. On land, some bacteria thrive in rocks as deep as 4 kilometers (2.5 miles) and in boiling hot springs. Moreover, air

many kilometers into the atmosphere. But even when we consider these extremes, life still must be thought of as being confined to a narrow band very near Earth's surface.

Plants and animals depend on the physical environment for the basics of life. However, organisms do not just respond to their physical environment. Through countless interactions, life-forms help maintain and alter the physical environment. Without life, the makeup and nature of the geosphere, hydrosphere, and atmosphere would be very different.

Geosphere

Lying beneath the atmosphere and the oceans is the solid Earth, or **geosphere**. The geosphere extends from the surface to the center of the planet, a depth of nearly 6400 kilometers (nearly 4000 miles), making it by far the largest of Earth's four spheres. Much of our study of the solid Earth focuses on the more accessible surface features. Fortunately, many of these features represent the outward expressions of the dynamic behavior of Earth's interior. By examining the most prominent surface features and their global extent, we can obtain clues to the dynamic processes that have shaped our planet. A first look at the structure of Earth's interior

Did You Know?

Primitive life first appeared in the oceans about 4 billion years ago and has been spreading and diversifying ever since.

currents can carry microorganisms

The ocean contains a significant portion of Earth's biosphere. Modern coral reefs are unique and complex examples and are home to about 25% of all marine species. Because of this diversity they are sometimes referred to as the ocean equivalent of a rain forest.

Figure 1.13 The biosphere The biosphere, one of Earth's four spheres, includes all life. (Coral reef photo by Darryl Leniuk/AGE Fotostock; rain forest photo by AGE Fotostock/SuperStock)

Tropical rain forests are characterized by hundreds of different species per square kilometer.

and at the major surface features of the geosphere will come later in the chapter.

Soil, the thin veneer of material at Earth's surface that supports the growth of plants, may be thought of as part of all four spheres. The solid portion is a mixture of weathered rock debris (geosphere) and organic matter from decayed plant and animal life (biosphere). The decomposed and disintegrated rock debris is the product of weathering processes that require air (atmosphere) and water (hydrosphere). Air and water also occupy the open spaces between the solid particles.

Concept Checks 1.4

- 1 List and describe Earth's four spheres.
- 2 Compare the height of the atmosphere to the thickness of the geosphere.
- (3) How much of Earth's surface do oceans cover? What percentage of Earth's total water supply do oceans represent?
- 4 To which sphere does soil belong?

1.5 Earth as a System

Define system and explain why Earth is considered to be a system.

Anyone who studies Earth soon learns that our planet is a dynamic body with many separate but interacting parts, or *spheres*. The hydrosphere, atmosphere, biosphere, and geosphere and all of their components can be studied separately. However, the parts are *not* isolated. Each is related in some way to the others, producing a complex and continuously interacting whole that we call the *Earth system*.

Earth System Science

A simple example of the interactions among different parts of the Earth system occurs every winter, as moisture evaporates from the Pacific Ocean and subsequently falls as rain in the hills and mountains of southern California, triggering destructive debris flows. The processes that move water from the hydrosphere to the atmosphere and then to the solid Earth have a profound impact on the plants and animals (including humans) that inhabit the affected regions (Figure 1.14).

Scientists have recognized that in order to more fully understand our planet, they must learn how its individual components (land, water, air, and life-forms) are interconnected. This endeavor, called **Earth system science**, aims to study Earth as a *system* composed of numerous interacting parts, or *subsystems*. Rather than look through the limited lens of only one of the traditional sciences geology, atmospheric science, chemistry, biology, and so on—Earth system science attempts to integrate the knowledge of several academic fields. Using an interdisciplinary approach, those engaged in Earth system science attempt to achieve the level of understanding necessary to comprehend and solve many of our global environmental problems.

A **system** is a group of interacting, or interdependent, parts that form a complex whole. Most of us hear and use the term *system* frequently. We may service our car's cooling *system*, make use of the city's transportation *system*, and be a participant in the political *system*. A news report might inform us of an approaching weather *system*. Further, we

know that Earth is just a small part of a larger system known as the *solar system*, which in turn is a subsystem of an even larger system called the Milky Way Galaxy.

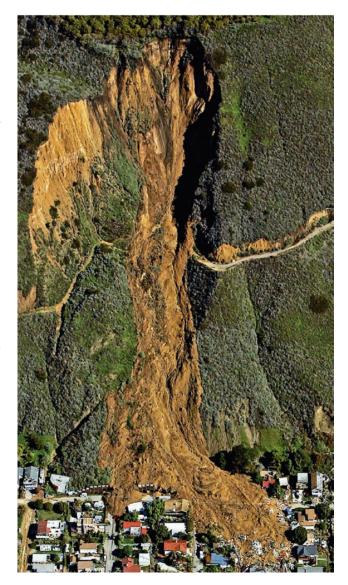


Figure 1.14 Deadly debris

flow This image provides an example of interactions among different parts of the Earth system. On January 10, 2005, extraordinary rains triggered this debris flow (popularly called a mudslide) in the coastal community of La Conchita, California. (AP Wideworld Photo) An Introduction to Geology

Did You Know?

Since 1970, Earth's average surface temperature has increased by about 0.6°C (1°F). By the end of the twenty-first century, the average global temperature may increase by an additional 2° to 4.5°C (3.5° to 8.1°F).

Figure 1.15 Change is a geologic constant When Mount St. Helens erupted in May 1980, the area shown here was buried by a volcanic mudflow. Now plants are reestablished, and new soil is forming. (Photo by Terry Donnelly/

Alamy Images)

The Earth System

The Earth system has a nearly endless array of subsystems in which matter is recycled over and over. One familiar loop or subsystem is the *hydrologic cycle*. It represents the unending circulation of Earth's water among the hydrosphere, atmosphere, biosphere, and geosphere. Water enters the atmosphere through evaporation from Earth's surface and transpiration from plants. Water vapor condenses in the atmosphere to form clouds, which in turn produce precipitation that falls back to Earth's surface. Some of the rain that falls onto the land sinks in to be taken up by plants or become groundwater, and some flows across the surface toward the ocean.

Viewed over long time spans, the rocks of the geosphere are constantly forming, changing, and re-forming. The loop that involves the processes by which one rock changes to another is called the *rock cycle* and will be discussed at some length later in the chapter. The cycles of the Earth system are not independent of one another. To the contrary, there are many places where the cycles come in contact and interact.

The parts of the Earth system are linked so that a change in one part can produce changes in any or all of the other parts. For example, when a volcano erupts, lava from Earth's interior may flow out at the surface and block a nearby valley. This new obstruction influences the region's drainage system by creating a lake or causing streams to change course. The large quantities of volcanic ash and gases that can be emitted during an eruption might be blown high into the atmosphere and influence the amount of solar energy that can reach Earth's surface. The result could be a drop in air temperatures over the entire hemisphere.



Where the surface is covered by lava flows or a thick layer of volcanic ash, existing soils are buried. This causes the soil-forming processes to begin anew to transform the new surface material into soil (Figure 1.15). The soil that eventually forms will reflect the interactions among many parts of the Earth system—the volcanic parent material, the climate, and the impact of biological activity. Of course, there would also be significant changes in the biosphere. Some organisms and their habitats would be eliminated by the lava and ash, whereas new settings for life, such as a lake formed by a lava dam, would be created. The potential climate change could also impact sensitive life-forms.

The Earth system is characterized by processes that vary on spatial scales from fractions of millimeters to thousands of kilometers. Time scales for Earth's processes range from milliseconds to billions of years. As we learn about Earth, it becomes increasingly clear that despite significant separations in distance or time, many processes are connected, and a change in one component can influence the entire system.

The Earth system is powered by energy from two sources. The Sun drives external processes that occur in the atmosphere, in the hydrosphere, and at Earth's surface. Weather and climate, ocean circulation, and erosional processes are driven by energy from the Sun. Earth's interior is the second source of energy. Heat remaining from when our planet formed and heat that is continuously generated by radioactive decay power the internal processes that produce volcanoes, earthquakes, and mountains.

Humans are *part of* the Earth system, a system in which the living and nonliving components are entwined and interconnected. Therefore, our actions produce changes in all the other parts. When we burn gasoline and coal, dispose of our wastes, and clear the land, we cause other parts of the system to respond, often in unforeseen ways. Throughout this book you will learn about many of Earth's subsystems, including the hydrologic system, the tectonic (mountain-building) system, the rock cycle, and the climate system. Remember that these components *and we humans* are all part of the complex interacting whole we call the Earth system.

Concept Checks 1.5

- 1) What is a system? List three examples.
- (2) What are the two sources of energy for the Earth system?
- 3 Predict how a change in the hydrologic cycle, such as increased rainfall in an area, might influence the biosphere and geosphere in that area.

1.6 Early Evolution of Earth

Outline the stages in the formation of our solar system.

Recent earthquakes caused by displacements of Earth's crust and lavas spewed from active volcanoes represent only the latest in a long line of events by which our planet has attained its present form and structure. The geologic processes operating in Earth's interior can be best understood when viewed in the context of much earlier events in Earth history.

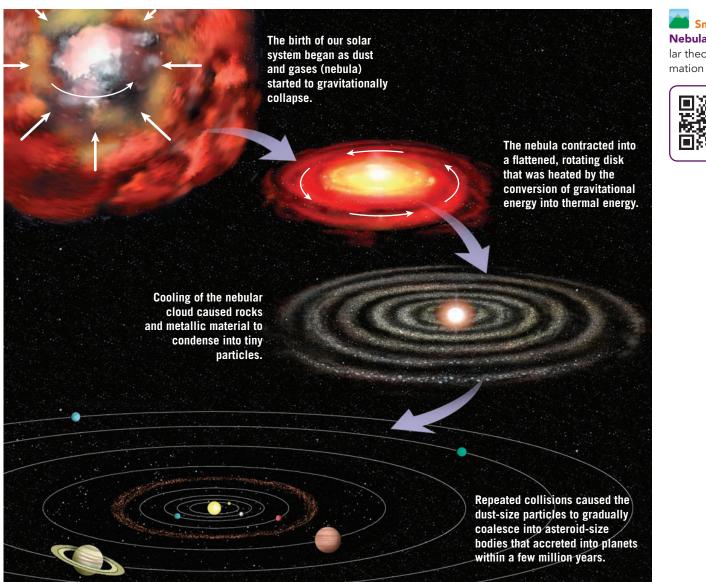
Origin of Planet Earth

This section describes the most widely accepted views on the origin of our solar system. The theory described here represents the most consistent set of ideas we have to explain what we know about our solar system today.

The Universe Begins Our scenario begins about 13.7 billion years ago, with the *Big Bang*, an incompre-

hensibly large explosion that sent all matter of the universe flying outward at incredible speeds. In time, the debris from this explosion, which was almost entirely hydrogen and helium, began to cool and condense into the first stars and galaxies. It was in one of these galaxies, the Milky Way, that our solar system and planet Earth took form.

The Solar System Forms Earth is one of eight planets that, along with several dozen moons and numerous smaller bodies, revolve around the Sun. The orderly nature of our solar system leads most researchers to conclude that Earth and the other planets formed at essentially the same time and from the same primordial material as the Sun. The **nebular theory** proposes that the bodies of our solar system evolved from an enormous rotating cloud called the **solar nebula** (**Figure 1.16**). Besides the hydrogen and helium atoms generated during the Big Bang, the



SmartFigure 1.16 Nebular theory The nebular theory explains the formation of the solar system.



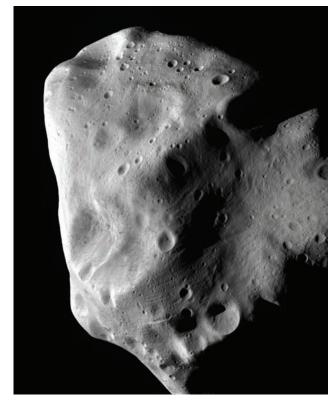


Figure 1.17 A remnant

planetesimal This image of Asteroid 21 Lutetia was obtained by special cameras aboard the *Rosetta* spacecraft on July 10, 2010. Spacecraft instruments showed that Lutetia is a primitive body (planetesimal) left over from when the solar system formed. (*Image courtesy of European Space Agency*) solar nebula consisted of microscopic dust grains and the ejected matter of long-dead stars. (Nuclear fusion in stars converts hydrogen and helium into the other elements found in the universe.)

Nearly 5 billion years ago, this huge cloud of gases and minute grains of heavier elements began to slowly contract due to the gravitational interactions among its particles. Some external influence, such as a shock wave traveling from a catastrophic explosion (supernova), may have triggered the collapse. As this slowly spiraling nebula contracted, it rotated faster and faster for the same reason ice skaters do when they draw their

arms toward their bodies. Eventually the inward pull of gravity came into balance with the outward force caused by the rotational motion of the nebula (see Figure 1.16). By this time, the once vast cloud had assumed a flat disk shape with a large concentration of material at its center called the *protosun* (pre-Sun). (Astronomers are fairly confident that the nebular cloud formed a disk because similar structures have been detected around other stars.)

During the collapse, gravitational energy was converted to thermal energy (heat), causing the temperature of the inner portion of the nebula to dramatically rise. At these high temperatures, the dust grains broke up into molecules and extremely energetic atomic particles. However, at distances beyond the orbit of Mars, the temperatures probably remained quite low. At -200° C (-328° F), the tiny particles in the outer portion of the nebula were likely covered with a thick layer of ices made of frozen water, carbon dioxide, ammonia, and methane. (Some of this material still resides in the outermost reaches of the solar system, in a region called the *Oort cloud*.) The disk-shaped cloud also contained appreciable amounts of the lighter gases hydrogen and helium.

The Inner Planets Form The formation of the Sun marked the end of the period of contraction and thus the end of gravitational heating. Temperatures in the region where the inner planets now reside began to decline. The decrease in temperature caused those substances with high melting points to condense into tiny particles that began to coalesce (join together). Materials such as iron and nickel and the elements of which the rock-forming minerals are composed—silicon, calcium, sodium, and so forth—formed metallic and rocky clumps that orbited the Sun (see

Figure 1.16). Repeated collisions caused these masses to coalesce into larger asteroid-size bodies, called *planetesimals*, which in a few tens of millions of years accreted into the four inner planets we call Mercury, Venus, Earth, and Mars (**Figure 1.17**). Not all of these clumps of matter were incorporated into the planetesimals. Those rocky and metallic pieces that remained in orbit are called *meteorites* when they survive an impact with Earth.

As more and more material was swept up by the planets, the high-velocity impact of nebular debris caused the temperature of these bodies to rise. Because of their relatively high temperatures and weak gravitational fields, the inner planets were unable to accumulate much of the lighter components of the nebular cloud. The lightest of these, hydrogen and helium, were eventually whisked from the inner solar system by the solar wind.

The Outer Planets Develop At the same time that the inner planets were forming, the larger, outer planets (Jupiter, Saturn, Uranus, and Neptune), along with their extensive satellite systems, were also developing. Because of low temperatures far from the Sun, the material from which these planets formed contained a high percentage of ices—water, carbon dioxide, ammonia, and methane—as well as rocky and metallic debris. The accumulation of ices accounts, in part, for the large size and low density of the outer planets. The two most massive planets, Jupiter and Saturn, had a surface gravity sufficient to attract and hold large quantities of even the lightest elements—hydrogen and helium.

Formation of Earth's Layered Structure

As material accumulated to form Earth (and for a short period afterward), the high-velocity impact of nebular debris and the decay of radioactive elements caused the temperature of our planet to steadily increase. During this time of intense heating, Earth became hot enough that iron and nickel began to melt. Melting produced liquid blobs of dense metal that sank toward the center of the planet. This process occurred rapidly on the scale of geologic time and produced Earth's dense iron-rich core.

Chemical Differentiation and Earth's Layers

The early period of heating resulted in another process of chemical differentiation, whereby melting formed buoyant masses of molten rock that rose toward the surface, where they solidified to produce a primitive crust. These rocky materials were enriched in oxygen and "oxygen-seeking" elements, particularly silicon and aluminum, along with lesser amounts of calcium, sodium, potassium, iron, and magnesium. In addition, some heavy metals such as gold, lead, and uranium, which have low melting points or were highly soluble in the ascending molten masses, were scavenged from Earth's interior and concentrated in the developing crust. This early period of chemical differentiation established the three basic divisions of Earth's interior the iron-rich *core*; the thin *primitive crust*; and Earth's largest layer, called the *mantle*, which is located between the core and crust.

An Atmosphere Develops An important consequence of the early period of chemical differentiation is that large quantities of gaseous materials were allowed to escape from Earth's interior, as happens today during volcanic eruptions. By this process, a primitive atmosphere gradually evolved. It is on this planet, with this atmosphere, that life as we know it came into existence.

Continents and Ocean Basins Evolve Following the events that established Earth's basic structure, the primitive crust was lost to erosion and other geologic processes, so we have no direct record of its makeup. When and exactly how the continental crust—and thus Earth's first landmasses—came into existence is a matter of ongoing research. Nevertheless, there is general agreement that the continental

crust formed gradually over the past 4 billion years. (The oldest rocks yet discovered are isolated fragments found in the Northwest Territories of Canada that have radiometric dates of about 4 billion years.) In addition, as you will see in subsequent chapters, Earth is an evolving planet whose continents and ocean basins have continually changed shape and even location during much of this period.

Concept Checks 1.6

- (1) Name and briefly outline the theory that describes the formation of our solar system.
- (2) List the inner planets and outer planets. Describe basic differences in size and composition.
- (3) Explain why density and buoyancy were important in the development of Earth's layered structure.

Did You Know?

The *light-year* is a unit for measuring distances to stars. Such distances are so large that familiar units such as kilometers or miles are cumbersome to use. One light-year is the distance light travels in one Earth year—about 9.5 trillion km (5.8 trillion mi)!

1.7 Earth's Internal Structure

Describe Earth's internal structure.

In the preceding section, you learned that the differentiation of material that began early in Earth's history resulted in the formation of three major layers defined by their chemical composition—the crust, mantle, and core. In addition to these compositionally distinct layers, Earth is divided into layers based on physical properties. The physical properties used to define such zones include whether the layer is solid or liquid and how weak or strong it is. Important examples include the lithosphere, asthenosphere, outer core, and inner core. Knowledge of both types of layers is important to our understanding of many geologic processes, including volcanism, earthquakes, and mountain building. **Figure 1.18** shows different views of Earth's layered structure.

How did we learn about the composition and structure of Earth's interior? We have never sampled the mantle or core directly. The nature of Earth's interior is determined by analyzing seismic waves from earthquakes. As these waves of energy penetrate the planet, they change speed and are bent and reflected as they move through zones that have different properties. Monitoring stations around the world detect and record this energy. With the aid of computers, these data are analyzed and used to build a detailed picture of Earth's interior. There is more about this in Chapter 9.

Earth's Crust

The **crust**, Earth's relatively thin, rocky outer skin, is of two different types—continental crust and oceanic crust. Both share the word *crust*, but the similarity ends there. The oceanic crust is roughly 7 kilometers (5 miles) thick and composed of the dark igneous rock *basalt*. By contrast, the continental crust averages about 35 kilometers (22 miles) thick but may exceed 70 kilometers (40 miles) in some mountainous regions such as the Rockies and

Himalayas. Unlike the oceanic crust, which has a relatively homogeneous chemical composition, the continental crust consists of many rock types. Although the upper crust has an average composition of a *granitic rock* called *granodiorite*, it varies considerably from place to place.

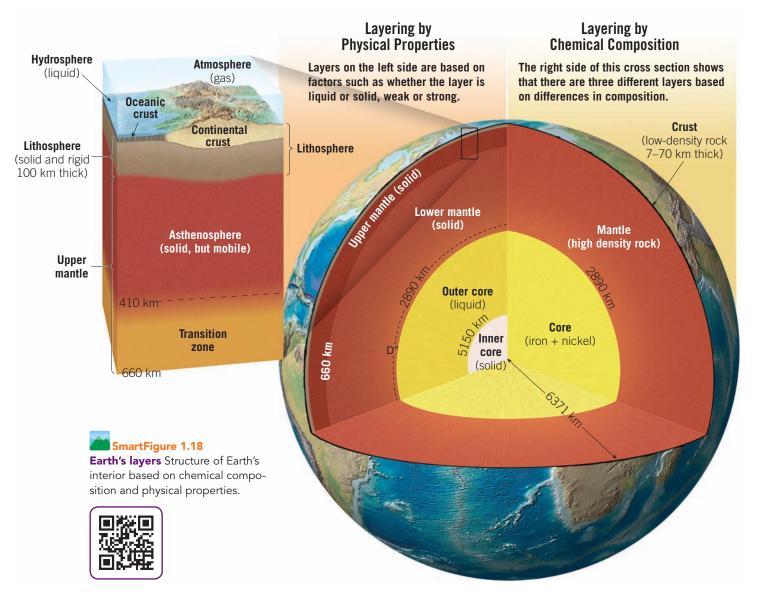
Continental rocks have an average density of about 2.7 g/cm³, and some have been discovered that are more than 4 billion years old. The rocks of the oceanic crust are younger (180 million years or less) and denser (about 3.0 g/cm³) than continental rocks. For comparison, liquid water has a density of 1 g/cm³; therefore, the density of basalt, the primary rock composing oceanic crust, is three times that of water.

Earth's Mantle

More than 82 percent of Earth's volume is contained in the **mantle**, a solid, rocky shell that extends to a depth of about 2900 kilometers (1800 miles). The boundary between the crust and mantle represents a marked change in chemical composition. The dominant rock type in the uppermost mantle is *peridotite*, which is richer in the metals magnesium and iron than the minerals found in either the continental or oceanic crust.

The Upper Mantle The upper mantle extends from the crust–mantle boundary down to a depth of about 660 kilometers (410 miles). The upper mantle can be divided into three different parts. The top portion of the upper mantle is part of the stronger *lithosphere*, and beneath that is the weaker *asthenosphere*. The bottom part of the upper mantle is called the *transition zone*.

The **lithosphere** (sphere of rock) consists of the entire crust plus the uppermost mantle and forms Earth's relatively cool, rigid outer shell (see Figure 1.18).



Did You Know?

We have never sampled the mantle or core directly. The structure of Earth's interior is determined by analyzing seismic waves from earthquakes. As these waves of energy penetrate Earth's interior, they change speed and are bent and reflected as they move through zones having different properties. Monitoring stations around the world detect and record this energy. Averaging about 100 kilometers (60 miles) thick, the lithosphere is more than 250 kilometers (155 miles) thick below the oldest portions of the continents. Beneath this stiff layer to a depth of about 410 kilometers (255 miles) lies a soft, comparatively weak layer known as the **asthenosphere** ("weak sphere"). The top portion of the asthenosphere has a temperature/pressure regime that results in a small amount of melting. Within this very weak zone, the lithosphere is mechanically detached from the layer below. The result is that the lithosphere is able to move independently of the asthenosphere, a fact we will consider in the next chapter.

It is important to emphasize that the strength of various Earth materials is a function of both their composition and the temperature and pressure of their environment. You should not get the idea that the entire lithosphere behaves like a rigid or brittle solid similar to rocks found on the surface. Rather, the rocks of the lithosphere get progressively hotter and weaker (more easily deformed) with increasing depth. At the depth of the uppermost asthenosphere, the rocks are close enough to their melting temperature (some melting may actually occur) that they are very easily deformed. Thus, the uppermost asthenosphere is weak because it is near its melting point, just as hot wax is weaker than cold wax.

From about 410 kilometers (255 miles) to about 660 kilometers (410 miles) in depth is the part of the upper mantle called the **transition zone**. The top of the transition zone is identified by a sudden increase in density from about 3.5 to 3.7 g/cm³. This change occurs because minerals in the rock peridotite respond to the increase in pressure by forming new minerals with closely packed atomic structures.

The Lower Mantle From a depth of 660 kilometers (410 miles) to the top of the core, at a depth of 2900 kilometers (1800 miles), is the **lower mantle**. Because of an increase in pressure (caused by the weight of the rock above), the mantle gradually strengthens with depth. Despite their strength, however, the rocks within the lower mantle are very hot and capable of extremely gradual flow.

In the bottom few hundred kilometers of the mantle is a highly variable and unusual layer called the D["] layer (pronounced "dee double-prime"). The nature of this boundary layer between the rocky mantle and the hot liquid iron outer core will be examined in Chapter 12.

Earth's Core

The **core** is thought to be composed of an iron–nickel alloy with minor amounts of oxygen, silicon, and sulfur—elements that readily form compounds with iron. At the extreme pressure found in the core, this iron-rich material has an average density of nearly 11 g/cm³ and approaches 14 times the density of water at Earth's center.

The core is divided into two regions that exhibit very different mechanical strengths. The **outer core** is

a *liquid layer* 2270 kilometers (1410 miles) thick. The movement of metallic iron within this zone generates Earth's magnetic field. The **inner core** is a sphere that has a radius of 1216 kilometers (754 miles). Despite its higher temperature, the iron in the inner core is *solid* due to the immense pressures that exist in the center of the planet.

Concept Checks 1.7

- List and describe the three major layers defined by their chemical composition.
- (2) Contrast the lithosphere and asthenosphere.
- 3 Distinguish between the outer core and the inner core.



Rocks and the Rock Cycle

Sketch, label, and explain the rock cycle.

Rock is the most common and abundant material on Earth. To a curious traveler, the variety seems nearly endless. When a rock is examined closely, we find that it usually consists of smaller crystals called minerals. *Minerals* are chemical compounds (or sometimes single elements), each with its own composition and physical properties. The grains or crystals may be microscopically small or easily seen with the unaided eye.

The minerals that compose a rock strongly influence its nature and appearance. In addition, a rock's *texture*—the size, shape, and/or arrangement of its constituent minerals—also has a significant effect on its appearance. A rock's mineral composition and texture, in turn, reflect the geologic processes that created it (**Figure 1.19**). Such analyses are critical to an understanding of our planet. This understanding has many practical applications, as in the search for energy and mineral resources and the solution of environmental problems.

Geologists divide rocks into three major groups: igneous, sedimentary, and metamorphic. **Figure 1.20** provides some examples. As you will learn, each group is linked to the others by the processes that act upon and within the planet.

Earlier in this chapter you learned that Earth is a system. This means that our planet consists of many interacting parts that form a complex whole. Nowhere is this idea better illustrated than when we examine the rock cycle (**Figure 1.21**). The **rock cycle** allows us to view many of the interrelationships among different parts of the Earth system. It helps us understand the origin of igneous, sedimentary, and metamorphic rocks and to see that each type is linked to the others by processes that act upon and within the planet. Consider the rock cycle to be a simplified but useful overview of physical geology. Learn the rock cycle well; you will be examining its interrelationships in greater detail throughout this book.

The Basic Cycle

Magma is molten rock that forms deep beneath Earth's surface. Over time, magma cools and solidifies. This process, called *crystallization*, may occur either beneath the surface or, following a volcanic eruption, at the surface. In either situation, the resulting rocks are called **igneous rocks**.

If igneous rocks are exposed at the surface, they undergo *weathering*, in which the day-in and day-out influences of the atmosphere slowly disintegrate and decompose rocks. The materials that result are often

Figure 1.19 Two basic

rock characteristics Texture and mineral composition are basic rock features. Both of these rock samples are about golf-ball size. (Basalt photo by Tyler Boyes/ Shutterstock; granite photo by geoz/Alamy Images)

The large crystals of light-colored minerals in granite result from the slow cooling of molten rock deep beneath the surface. Granite is abundant in the continental crust.

Basalt is rich in dark minerals. Rapid cooling of molten rock at Earth's surface is responsible for the rock's microscopically small crystals. Oceanic crust is composed mainly of basalt. moved downslope by gravity before being picked up and transported by any of a number of erosional agents, such as running water, glaciers, wind, or waves. Eventually these particles and dissolved substances, called sediment, are deposited. Although most sediment ultimately comes to rest in the ocean, other sites of deposition include river floodplains, desert basins, swamps, and sand dunes.

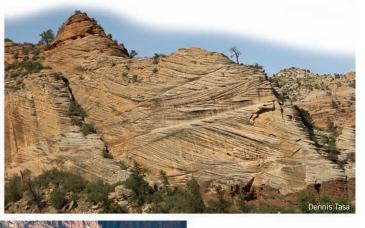
Next, the sediments undergo lithification, a term meaning "conversion into rock." Sediment is usually lithified into sedimentary rock when compacted by the weight of overlying layers or when cemented as percolating groundwater fills the pores with mineral matter.

If the resulting sedimentary rock is buried deep within igneous, sedimentary, and Earth and involved in the dynamics of mountain building or



Igneous rocks form when molten rock solidifies at the surface (extrusive) or beneath the surface (intrusive). The lava flow in the foreground is the fine-grained rock basalt and came from SP Crater in northern Arizona.

Sedimentary rocks consist of particles derived from the weathering of other rocks. This layer consists of durable sand-size grains of the glassy mineral quartz that are cemented into a solid rock. The grains were once a part of extensive dunes. This rock layer, called the Navajo Sandstone, is prominent in southern Utah.





The metamorphic rock pictured here, known as the Vishnu Schist, is exposed in the inner gorge of the Grand Canyon. Its formation is associated with environments deep below Earth's surface where temperatures and pressures are high and with the forces associated with ancient mountain-building processes that occurred in Precambrian time.

intruded by a mass of magma, it is subjected to great pressures and/or intense heat. The sedimentary rock reacts to the changing environment and turns into the third rock type, metamorphic rock. When metamorphic rock is subjected to additional pressure changes or to still higher temperatures, it melts, creating magma, which eventually crystallizes into igneous rock, starting the cycle all over again.

Where does the energy that drives Earth's rock cycle come from? Processes driven by heat from Earth's interior are responsible for creating igneous and metamorphic rocks. Weathering and erosion, external processes powered by energy from the Sun, produce the sediment from which sedimentary rocks form.

Alternative Paths

The paths shown in the basic cycle are not the only ones that are possible. To the contrary, other paths are just as likely to be followed as those described in the preceding section. These alternatives are indicated by the light blue arrows in Figure 1.21.

Rather than being exposed to weathering and erosion at Earth's surface, igneous rocks may remain deeply buried. Eventually these masses may be subjected to the strong compressional forces and high temperatures associated with mountain building. When this occurs, they are transformed directly into metamorphic rocks.

Metamorphic and sedimentary rocks, as well as sediment, do not always remain buried. Rather, overlying layers may be stripped away, exposing the once-buried rock. When this happens, the material is attacked by weathering processes and turned into new raw materials for sedimentary rocks.

Although rocks may seem to be unchanging masses, the rock cycle shows that they are not. The changes, however, take time—great amounts of time. We can observe different parts of the cycle operating all over the world. Today new magma is forming beneath the island of Hawaii. When it erupts at the surface, the lava flows add to the size of the island. Meanwhile, the Colorado Rockies are gradually being worn down by weathering and erosion. Some of this weathered debris will eventually be carried to the Gulf of Mexico, where it will add to the already substantial mass of sediment that has accumulated there.

Concept Checks 1.8

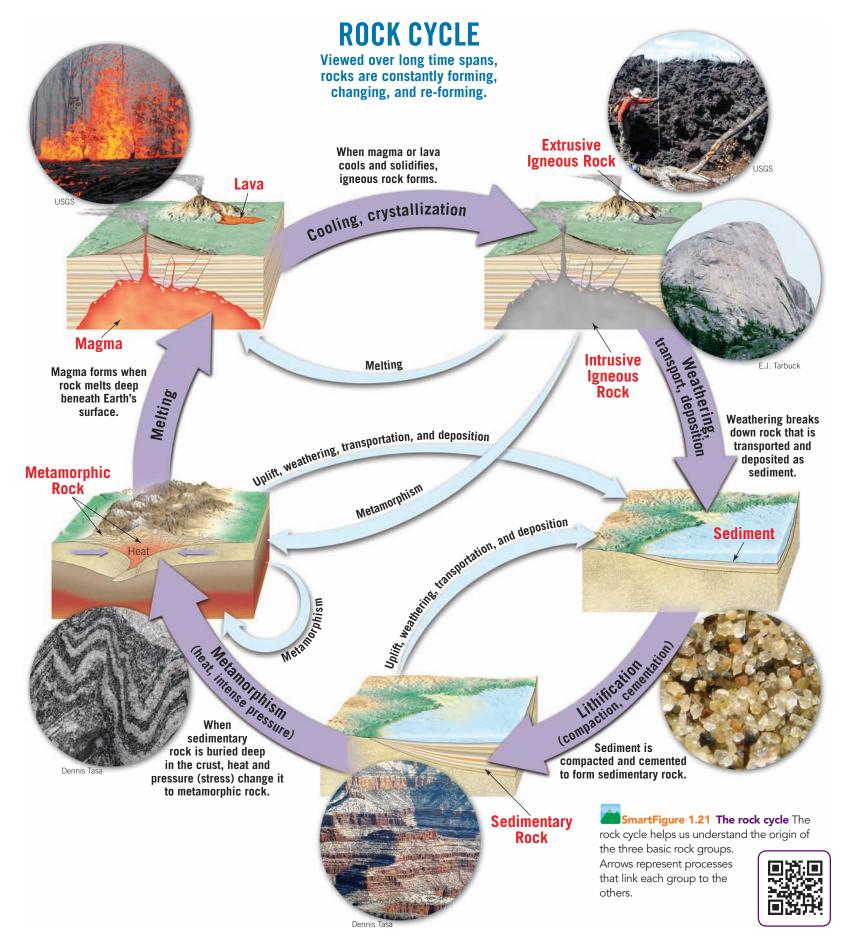
- List two rock characteristics that are used to determine the processes that created a rock.
- (2)Sketch and label a basic rock cycle. Make sure to include alternate paths.
- Use the rock cycle to explain the statement "One rock is the raw material for another."

Figure 1.20 Three rock

groups Geologists divide

rocks into three groups-

metamorphic.



42

1.9

The Face of Earth

List and describe the major features of the continents and ocean basins.

The two principal divisions of Earth's surface are the **continents** and the **ocean basins** (Figure 1.22). A significant difference between these two areas is their relative levels. The elevation difference between the continents

and ocean basins is primarily a result of differences in their respective densities and thicknesses:

• **Continents.** The continents are remarkably flat features that have the appearance of plateaus protruding above

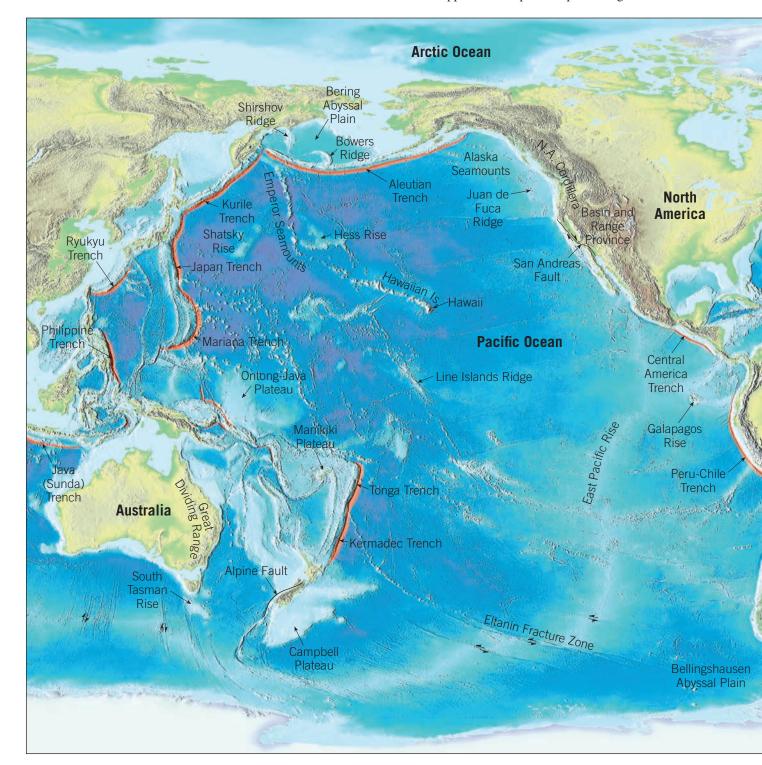
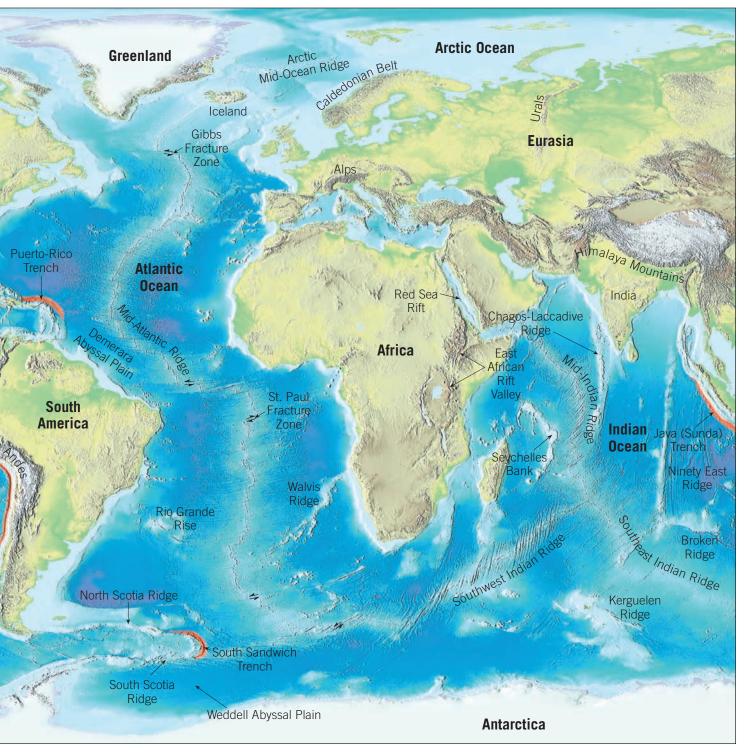


Figure 1.22 The face of Earth Major surface features of the geosphere. sea level. With an average elevation of about 0.8 kilometer (0.5 mile), continental blocks lie close to sea level, except for limited areas of mountainous terrain. Recall that the continents average about 35 kilometers (22 miles) thick and are composed of granitic rocks that have a density of about 2.7 g/cm³.

• Ocean basins. The average depth of the ocean floor is about 3.8 kilometers (2.4 miles) below sea level, or about 4.5 kilometers (2.8 miles) lower than the average elevation of the continents. The basaltic rocks that comprise the oceanic crust average only 7 kilometers (5 miles) thick and have an average density of about 3.0 g/cm³.

Thus, the thicker and less dense continental crust is more buoyant than the oceanic crust. As a result, continental crust floats on top of the deformable rocks of the mantle at a higher level than oceanic crust for the same reason that a large, empty (less dense) cargo ship rides higher than a small, loaded (denser) one.



Did You Know?

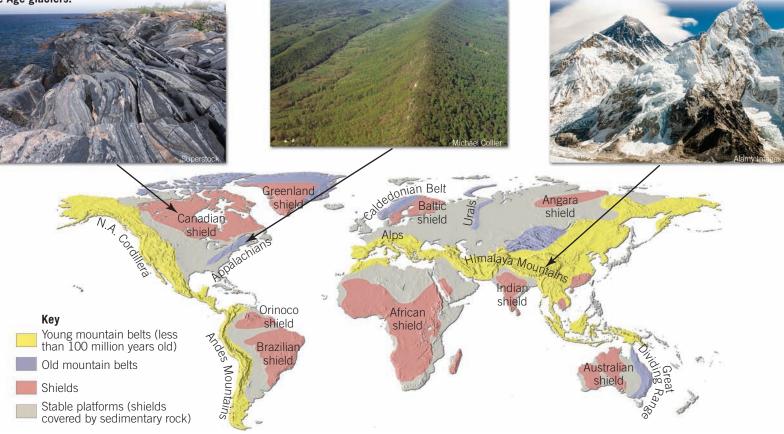
Ocean depths are often expressed in fathoms. One fathom equals 1.8 m or 6 ft, which is about the distance of a person's outstretched arms. The term is derived from how depth-sounding lines were brought back on board a vessel by hand. As the line was hauled in, a worker counted the number of arm lengths collected. By knowing the length of the person's outstretched arms, the amount of line taken in could be calculated. The length of 1 fathom was later standardized to 6 ft.

The Canadian Shield is an expansive region of ancient Precambrian rocks, some more than 4 billion years old. It was recently scoured by Ice Age glaciers.



The Appalachians are old mountains. Mountain building began about 480 million years ago and continued for more than 200 million years. Erosion has lowered these once lofty peaks.

The rugged Himalayas are the highest mountains on Earth and are geologically young. They began forming about 50 million years ago and uplift continues today.



martFigure 1.23

The continents Distribution of mountain belts, stable platforms, and shields.



Major Features of the Continents

The major features of the continents can be grouped into two distinct categories: uplifted regions of deformed rocks that make up present-day mountain belts and extensive flat, stable areas that have eroded nearly to sea level. Notice in Figure 1.23 that the young mountain belts tend to be long, narrow features at the margins of continents and that the flat, stable areas are typically located in the interior of the continents.

Mountain Belts The most prominent features of the continents are mountains. Although the distribution of mountains appears to be random, this is not the case. The youngest mountains (those less than 100 million years old) are located principally in two major zones. The circum-Pacific belt (the region surrounding the Pacific Ocean) includes the mountains of the western Americas and continues into the western Pacific, in the form of volcanic island arcs (see Figure 1.22). Island arcs are active mountainous regions composed largely of volcanic rocks and deformed sedimentary rocks. Examples include the Aleutian islands, Japan, the Philippines, and New Guinea.

The other major mountain belt extends eastward from the Alps through Iran and the Himalayas and then dips southward into Indonesia. Careful examination of mountainous terrains reveals that most are places where thick sequences of rocks have been squeezed and highly deformed, as if placed in a gigantic vise. Older mountains are also found on the continents. Examples include the Appalachians in the eastern United States and the Urals in Russia. Their once lofty peaks are now worn low, as a result of millions of years of weathering and erosion.

The Stable Interior Unlike the young mountain belts, which have formed within the past 100 million years, the interiors of the continents, called cratons, have been relatively stable (undisturbed) for the past 600 million years or even longer. Typically these regions were involved in mountain-building episodes much earlier in Earth's history.

Within the stable interiors are areas known as shields, which are expansive, flat regions composed largely of deformed igneous and metamorphic rocks. Notice in Figure 1.23 that the Canadian Shield is exposed in much of the northeastern part of North America. Radiometric dating of various shields has revealed that they are truly ancient regions. All contain Precambrian-age rocks that

are more than 1 billion years old, with some samples approaching 4 billion years in age. Even these oldestknown rocks exhibit evidence of enormous forces that have folded, faulted, and metamorphosed them. Thus, we conclude that these rocks were once part of an ancient mountain system that has since been eroded away to produce these expansive, flat regions.

Other flat areas of the craton exist, in which highly deformed rocks, like those found in the shields, are covered by a relatively thin veneer of sedimentary rocks. These areas are called **stable platforms**. The sedimentary rocks in stable platforms are nearly horizontal, except where they have been warped to form large basins or domes. In North America a major portion of the stable platform is located between the Canadian Shield and the Rocky Mountains.

Major Features of the Ocean Floor

If all water were drained from the ocean basins, a great variety of features would be seen, including chains of volcanoes, deep canyons, plateaus, and large expanses of monotonously flat plains. In fact, the scenery would be nearly as diverse as that on the continents (see Figure 1.22).

During the past 65 years, oceanographers have used modern depth-sounding equipment and satellite technology to map significant portions of the ocean floor. These studies have led them to identify three major regions: *continental margins, deep-ocean basins,* and *oceanic (mid-ocean) ridges.*

Continental Margin The **continental margin** is the portion of the seafloor adjacent to major landmasses. It may include the *continental shelf*, the *continental slope*, and the *continental rise*.

Although land and sea meet at the shoreline, this is *not* the boundary between the continents and the ocean basins. Rather, along most coasts, a gently sloping platform of material, called the **continental shelf**, extends seaward from the shore. Because it is underlain by continental crust, it is clearly a flooded extension of the continents. A glance at Figure 1.22 shows that the width of the continental shelf is variable. For example, it is broad along the east and Gulf coasts of the United States but relatively narrow along the Pacific margin of the continent.

The boundary between the continents and the deepocean basins lies along the **continental slope**, which is a relatively steep dropoff that extends from the outer edge of the continental shelf to the floor of the deep ocean (see Figure 1.22). Using this as the dividing line, we find that about 60 percent of Earth's surface is represented by ocean basins and the remaining 40 percent by continents.

In regions where trenches do not exist, the steep continental slope merges into a more gradual incline known as the **continental rise**. The continental rise consists of a thick wedge of sediment that moved downslope from the continental shelf and accumulated on the deep-ocean floor. **Deep-Ocean Basins** Between the continental margins and oceanic ridges lie the **deep-ocean basins**. Parts of these regions consist of incredibly flat features called **abyssal plains**. The ocean floor also contains extremely deep depressions that are occasionally more than 11,000 meters (36,000 feet) deep. Although these **deep-ocean trenches** are relatively narrow and represent only a small fraction of the ocean floor, they are nevertheless very significant features. Some trenches are located adjacent to young mountains that flank the continents. For example, in Figure 1.22 the Peru–Chile trench off the west coast of South America parallels the Andes Mountains. Other trenches parallel island chains called *volcanic island arcs*.

Dotting the ocean floor are submerged volcanic structures called **seamounts**, which sometimes form long, narrow chains. Volcanic activity has also produced several large *lava plateaus*, such as the Ontong Java Plateau located northeast of New Guinea. In addition, some submerged plateaus are composed of continental-type crust. Examples include the Campbell Plateau southeast of New Zealand and the Seychelles Bank northeast of Madagascar.

Oceanic Ridges The most prominent feature on the ocean floor is the **oceanic ridge**, or **mid-ocean ridge**. As shown in Figure 1.22, the Mid-Atlantic Ridge and the East Pacific Rise are parts of this system. This broad elevated feature forms a continuous belt that winds for more than 70,000 kilometers (43,000 miles) around the globe, in a manner similar to the seam of a baseball. Rather than consist of highly deformed rock, such as most of the mountains on the continents, the oceanic ridge system consists of layer upon layer of igneous rock that has been fractured and uplifted.

Being familiar with the topographic features that comprise the face of Earth is essential to understanding the mechanisms that have shaped our planet. What is the significance of the enormous ridge system that extends through all the world's oceans? What is the connection, if any, between young, active mountain belts and oceanic trenches? What forces crumple rocks to produce majestic mountain ranges? These are a few of the questions that will be addressed in the next chapter, as we begin to investigate the dynamic processes that shaped our planet in the geologic past and will continue to shape it in the future.

Concept Checks 1.9

- (1) Compare and contrast continents and ocean basins.
- 2 Describe the general distribution of Earth's youngest mountains.
- What is the difference between shields and stable platforms?
- What are the three major regions of the ocean floor, and what are some features associated with each?

oncepts in Review An Introduction to Geology

1.1 Geology: The Science of Earth

Distinguish between physical and historical geology and describe the connections between people and geology.

Key Terms: geology, physical geology, historical geology

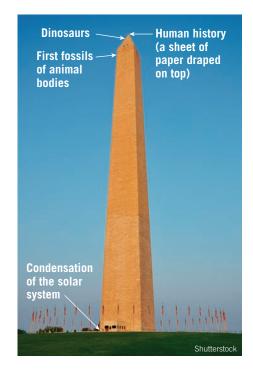
- Geologists study Earth. Physical geologists focus on the processes by which Earth operates and the materials that result from those processes. Historical geologists apply an understanding of Earth materials and processes to reconstruct the history of our planet.
- People have a relationship with planet Earth that can be positive and negative. Earth processes and products sustain us every day, but they can also harm us. Similarly, people have the ability to alter or harm natural systems, including those that sustain civilization.
 - Consider the question of when a given volcano is likely to erupt, and also the question of whether volcanic eruptions played a part in the extinction of the dinosaurs. Which is an issue that a physical geologist would address? Which question would a historical geologist focus on?

1.2 The Development of Geology

Summarize early and modern views on how change occurs on Earth and relate them to the prevailing ideas about the age of Earth.

Key Terms: catastrophism, uniformitarianism

- Early ideas about the nature of Earth were based on religious traditions and notions of great catastrophes. In 1795, James Hutton emphasized that the same slow processes have acted over great spans of time and are responsible for Earth's rocks, mountains, and landforms. This similarity of process over vast spans of time led to this principle being dubbed "uniformitarianism."
- Based on the rate of radioactive decay of certain elements, the age of Earth has been calculated to be about 4,600,000,000 (4.6 billion) years. That is an incredibly vast amount of time.
 - The Washington Monument in Washington, DC, is a little less than 169.294 meters (555.5 feet) tall. A sheet of paper is about 0.7 millimeter thick. If you divide 0.7 millimeter by 169,294 millimeters, the result is about 0.0004 percent of the total.



1.3 The Nature of Scientific Inquiry

Discuss the nature of scientific inquiry, including the construction of hypotheses and the development of theories.

Key Terms: hypothesis, theory

• Geologists make observations, construct tentative explanations for those observations (hypotheses), and then test those hypotheses with field investigations and laboratory work. In science, a theory is a well-tested and widely accepted view that the scientific community agrees best explains certain observable facts.

• As flawed hypotheses are discarded, scientific knowledge moves closer to a correct understanding, but we can never be fully confident that we know all the answers. Scientists must always be open to new information that forces a change in our model of the world.

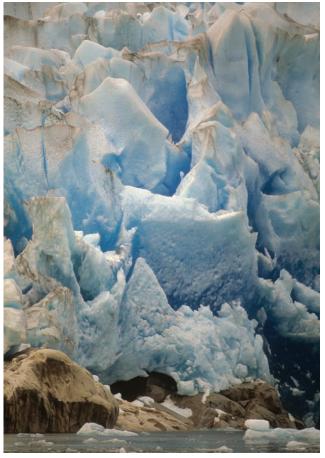
1.4 Earth's Spheres

List and describe Earth's four major spheres.

Key Terms: hydrosphere, atmosphere, biosphere, geosphere

- Earth's physical environment is traditionally divided into three major parts: the solid Earth, called the geosphere; the water portion of our planet, called the hydrosphere; and Earth's gaseous envelope, called the atmosphere.
- A fourth Earth sphere is the biosphere, the totality of life on Earth. It is concentrated in a relatively thin zone that extends a few kilometers into the hydrosphere and geosphere and a few kilometers up into the atmosphere.
- Of all the water on Earth, more than 96 percent is in the oceans, which cover nearly 71 percent of the planet's surface.

Is glacial ice part of the geosphere, or does it belong to the hydrosphere? Explain your answer.



1.5 Earth as a System

Define system and explain why Earth is considered to be a system.

Key Terms: Earth system science, system

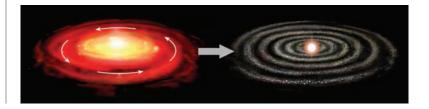
- Although each of Earth's four spheres can be studied separately, they are all related in a complex and continuously interacting whole that is called the Earth system.
- Earth system science uses an interdisciplinary approach to integrate the knowledge of several academic fields in the study of our planet and its global environmental problems.
- The two sources of energy that power the Earth system are (1) the Sun, which drives the external processes that occur in the atmosphere, hydrosphere, and at Earth's surface, and (2) heat from Earth's interior that powers the internal processes that produce volcanoes, earthquakes, and mountains.
 - Give a specific example of how humans are affected by the
 Earth system and another example of how humans affect the Earth system.

1.6 Early Evolution of Earth

Outline the stages in the formation of our solar system.

Key Terms: nebular theory, solar nebula

- The nebular theory describes the formation of the solar system. The planets and Sun began forming about 5 billion years ago from a large cloud of dust and gases.
- As the cloud contracted, it began to rotate and assume a disk shape. Material that was gravitationally pulled toward the center became the protosun. Within the rotating disk, small centers, called planetesimals, swept up more and more of the cloud's debris.
- Because of their high temperatures and weak gravitational fields, the inner planets were unable to accumulate and retain many of the lighter components. Because of the very cold temperatures existing far from the Sun, the large outer planets consist of huge amounts of lighter materials. These gaseous substances account for the comparatively large sizes and low densities of the outer planets.
 - Earth is about 4.6 billion years old. If all of the planets in our solar system formed at about the same time, how old would you expect Mars to be? Jupiter? The Sun?

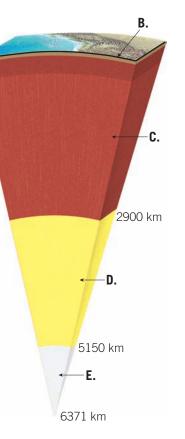


1.7 Earth's Internal Structure

Describe Earth's internal structure.

Key Terms: crust, mantle, lithosphere, asthenosphere, transition zone, lower mantle, core, outer core, inner core

- Compositionally, the solid Earth has three layers: core, mantle, and crust. The core is most dense, and the crust is least dense.
- Earth's interior can also be divided into layers based on physical properties. The crust and upper mantle make a two-part layer called the lithosphere, which is broken into the plates of plate tectonics. Beneath that is the "weak" asthenosphere. The lower mantle is stronger than the asthenosphere and overlies the molten outer core. This liquid is made of the same iron-nickel alloy as the inner core, but the extremely high pressure of Earth's center compacts the inner core into a solid form.



The diagram represents Earth's layered structure. Does it show layering based on physical properties or layering based on composition? Identify the lettered layers.

1.9 The Face of Earth

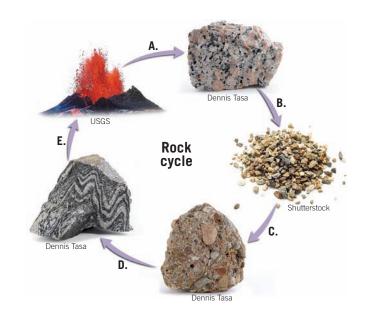
List and describe the major features of the continents and ocean basins.

- Key Terms: continent, ocean basin, mountain belt, craton, shield, stable platform, continental margin, continental shelf, continental slope, continental rise, deep-ocean basin, abyssal plain, deep-ocean trench, seamount, oceanic ridge (mid-ocean ridge)
- Two principal divisions of Earth's surface are the continents and ocean basins. A significant difference is their relative levels. The elevation differences between continents and ocean basins is primarily the result of differences in their respective densities and thicknesses.

1.8 Rocks and the Rock Cycle

Sketch, label, and explain the rock cycle.

- Key Terms: rock cycle, igneous rock, sediment, sedimentary rock, metamorphic rock
- The rock cycle is a good model for thinking about the transformation of one rock to another due to Earth processes. All igneous rocks are made from molten rock. All sedimentary rocks are made from weathered products of other rocks. All metamorphic rocks are the products of preexisting rocks that are transformed at high temperatures or pressures. Given the right conditions, any kind of rock can be transformed into any other kind of rock.
 - Name the processes that are represented by each of the letters in this simplified rock cycle diagram.



- Continents consist of relatively flat, stable areas called cratons. Where a craton is blanketed by a relatively thin layer of sediment or sedimentary rock, it is called a stable platform. Where a craton is exposed at the surface, it is known as a shield. Wrapping around the edges of some cratons are mountain belts, linear zones of intense deformation and metamorphism.
- There are shallow portions of the oceans that are essentially flooded margins of the continents, and there are deeper portions that include vast abyssal plains and deep ocean trenches. Seamounts and lava plateaus interrupt the abyssal plain in some places.
 - Put these features of the ocean floor in order from shallowest to deepest: continental slope, deep-ocean trench, continental shelf, abyssal plain, continental rise.

Give It Some Thought

- (1) The geologic time scale represents the relative chronology as recorded in rocks, exposed on the Earth's surface. The time scale does not give us an idea of absolute age of the rocks and the date of events. How can the absolute age of the rocks be measured? What is the difference between age determination of a rock and dating an event?
- 2 After entering a dark room, you turn on a wall switch, but the light does not come on. Suggest at least three hypotheses that might explain this observation. Once you have formulated your hypotheses, what is the next logical step?



- 3 Based on what is discussed in this Chapter answer the following questions.
 - a. Earth system science attempts to integrate the knowledge of several academic fields. Why is the integrated approach necessary?
 - **b.** How do human beings contribute to the natural geologic processes operating in the earth system? Is modernization and development of the human society posing a threat to natural geological processes?
 - **c.** Which mechanisms, internal or external, combine to allow Earth's temperature to favor life?
- 4 The Himalayas, which were formed almost 50 million years ago, are considered to be geologically young. Due to tectonic adjustments they are still rising. What could be the effect of this rising landform? What could be the probable effects on the Ganges-Brahmaputra delta?
- (5) This jet is cruising at an altitude of 10 kilometers (6.2 miles). Refer to the graph in Figure 1.12. What is the approximate air pressure at the altitude where the jet is flying? About what percentage of the atmosphere is below the jet (assuming that the pressure at the surface is 1000 millibars)?



6 The accompanying photo provides an example of interactions among different parts of the Earth system. It is a view of a mudflow that was triggered by extraordinary rains. Which of Earth's four

spheres were involved in this natural disaster that buried a small town on the Philippine island of Leyte? Describe how each contributed to or was influenced by the event.



- 7 Given the geological principle that the present is the key to the past, also known as Uniformitarianism, can we use this in reverse, i.e. use evidence of processes recorded in geological records to interpret contemporary geological events?
- (8) This photo shows the picturesque coastal bluffs and rocky shoreline along a portion of the California coast south of San Simeon State Park. This area, like other shorelines, is described as an *interface*. What does this mean? Does the shoreline represent the boundary between the continent and ocean basin? Explain.



Plate Tectonics: A Scientific Revolution

2



ocus on Concepts

Each statement represents the primary learning objective for the corresponding major heading within the chapter. After you complete the chapter, you should be able to:

- 2.1 Discuss the view that most geologists held prior to the 1960s regarding the geographic positions of the ocean basins and continents.
- **2.2** List and explain the evidence Wegener presented to support his continental drift hypothesis.
- **2.3** Discuss the two main objections to the continental drift hypothesis.
- 2.4 List the major differences between Earth's lithosphere and asthenosphere and explain the importance of each in the plate tectonics theory.
- **2.5** Sketch and describe the movement along a divergent plate boundary that results in the formation of new oceanic lithosphere.
- **2.6** Compare and contrast the three types of convergent plate boundaries and name a location where each type can be found.
- 2.7 Describe the relative motion along a transform plate boundary and locate several examples on a plate boundary map.
- **2.8** Explain why plates such as the African and Antarctic plates are getting larger, while the Pacific plate is getting smaller.
- **2.9** List and explain the evidence used to support the plate tectonics theory.
- **2.10** Describe two methods researchers use to measure relative plate motion.
- 2.11 Summarize what is meant by platemantle convection and explain two of the primary driving forces of plate motion.

This climber is ascending Mount Paterno in the Dolomites, a mountain range in northeastern Italy that is part of the Alps. (Photo by allesfoto/imagebroker/AGE Fotostock) **Plate tectonics is the first theory** to provide a comprehensive view of the processes that produced Earth's major surface features, including the continents and ocean basins. Within the framework of this theory, geologists have found explanations for the basic causes and distribution of earthquakes, volcanoes, and mountain belts. Further, we are now better able to explain the distribution of plants and animals in the geologic past, as well as the distribution of economically significant mineral deposits.

Did You Know?

Although Alfred Wegener is rightfully credited with formulating the continental drift hypothesis, he was not the first to suggest continental mobility. An American geologist, F. B. Taylor, published the first paper to outline this important idea. However, Taylor's paper provided little supporting evidence, whereas Wegener spent much of his professional life trying to substantiate his views.

Figure 2.1 Rock pinnacle near Mount Blanc The Alps were created by the collision of the African and Eurasian plates. (Photo by Bildagentur Walhaeus/AGE Fotostock)

2.1 From Continental Drift to Plate Tectonics

Discuss the view that most geologists held prior to the 1960s regarding the geographic positions of the ocean basins and continents.

Prior to the late 1960s, most geologists held the view that the ocean basins and continents had fixed geographic positions and were of great antiquity. Researchers came to realize that Earth's continents are not static; instead, they gradually migrate across the globe. Because of these movements, blocks of continental material collide, deforming the intervening crust, thereby creating Earth's great mountain chains (**Figure 2.1**). Furthermore, landmasses occasionally split apart. As continental blocks separate, a new ocean basin emerges between them. Meanwhile, other portions of the seafloor plunge into the mantle. In short, a dramatically different model of Earth's tectonic processes emerged. Tectonic processes are processes that deform Earth's crust to create major structural features, such as mountains, continents, and ocean basins.

This profound reversal in scientific thought has been appropriately described as a *scientific revolution*. The revolution began early in the twentieth century, as a relatively straightforward proposal called *continental drift*. For more than 50 years, the scientific establishment categorically rejected the idea that continents are capable of movement. Continental drift was particularly distasteful to North American geologists, perhaps because much of the supporting evidence had been gathered from the continents of Africa, South America, and Australia, with which most North American geologists were unfamiliar.



Following World War II, modern instruments replaced rock hammers as the tools of choice for many researchers. Armed with more advanced tools, geologists and a new breed of researchers, including *geophysicists* and *geochemists*, made several surprising discoveries that began to rekindle interest in the drift hypothesis. By 1968 these developments had led to the unfolding of a far more encompassing explanation known as the *theory of plate tectonics*.

In this chapter, we will examine the events that led to this dramatic reversal of scientific opinion. We will also briefly trace the development of the *continental drift* *hypothesis*, examine why it was initially rejected, and consider the evidence that finally led to the acceptance of its direct descendant—the theory of plate tectonics.

Concept Checks 2.1

- (1) Briefly describe the view held by most geologists regarding the ocean basins and continents prior to the 1960s.
- 2 What group of geologists were the least receptive to the continental drift hypothesis? Explain.

2.2 Continental Drift: An Idea Before Its Time

List and explain the evidence Wegener presented to support his continental drift hypothesis.

The idea that continents, particularly South America and Africa, fit together like pieces of a jigsaw puzzle came about during the 1600s, as better world maps became available. However, little significance was given to this notion until 1915, when Alfred Wegener (1880–1930), a German meteorologist and geophysicist, wrote *The Origin of Continents and Oceans*. This book set forth the basic outline of Wegener's hypothesis, called **continental drift**, which dared to challenge the long-held assumption that the continents and ocean basins had fixed geographic positions.

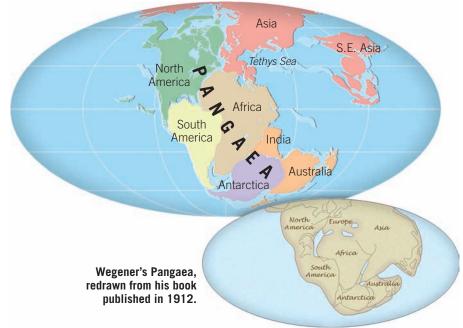
Wegener suggested that a single **supercontinent** consisting of all Earth's landmasses once existed.* He named this giant landmass **Pangaea** (pronounced "Pan-jee-ah," meaning "all lands") (**Figure 2.2**). Wegener further hypothesized that about 200 million years ago, during the early part of the Mesozoic era, this supercontinent began to fragment into smaller landmasses. These continental blocks then "drifted" to their present positions over a span of millions of years.

Wegener and others who advocated the continental drift hypothesis collected substantial evidence to support their point of view. The fit of South America and Africa and the geographic distribution of fossils and ancient climates seemed to buttress the idea that these now separate landmasses were once joined. Let us examine some of this evidence.

Evidence: The Continental Jigsaw Puzzle

Like a few others before him, Wegener suspected that the continents might once have been joined when he noticed

Modern reconstruction of Pangaea



the remarkable similarity between the coastlines on opposite sides of the Atlantic Ocean. However, other Earth scientists challenged Wegener's use of present-day shorelines to fit these continents together. These opponents correctly argued that shorelines are continually modified by wave erosion and depositional processes. Even if continental displacement had taken place, a good fit today would be unlikely. Because Wegener's original jigsaw fit of the continents was crude, it is assumed that he was aware of this problem (see Figure 2.2).

Scientists later determined that a much better approximation of the outer boundary of a continent is the seaward edge of its continental shelf, which lies submerged a few

SmartFigure 2.2 Reconstructions of Pangaea The supercontinent of Pangaea, as it is thought to have appeared 200 million years ago.



^{*}Wegener was not the first person to conceive of a long-vanished supercontinent. Edward Suess (1831–1914), a distinguished nineteenth-century geologist, pieced together evidence for a giant landmass consisting of the continents of South America, Africa, India, and Australia.

A F R I C A



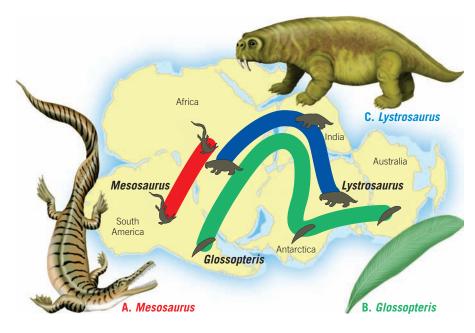
Figure 2.3 Two of the puzzle pieces The best fit of South America and Africa is along the continental slope at a depth of 500 fathoms (about 900 meters [3000 feet]). (Based on A. G. Smith, "Continental Drift," in Understanding the Earth, edited by I. G. Gass, Artemis Press.)

hundred meters below sea level. In the early 1960s, Sir Edward Bullard and two associates constructed a map that pieced together the edges of the continental shelves of South America and Africa at a depth of about 900 meters (3000 feet) (Figure 2.3). The remarkable fit that was obtained was more precise than even these researchers had expected.

Evidence: Fossils Matching Across the Seas

Although the seed for Wegener's hypothesis came from the remarkable similarities of the continental margins on opposite sides of the Atlantic, it was when he learned that identical fossil organisms had been discovered in rocks from both South America and Africa that his pursuit of continental drift became more focused. Through a review of the literature, Wegener learned that most paleontologists (scientists who study the fossilized remains of ancient organisms) were in agreement that some type of land connection was needed to explain the existence of similar Mesozoic age life-forms on widely separated landmasses. Just as modern life-forms native to North America are quite different from those of Africa and Australia, during the Mesozoic era, organisms on widely separated continents should have been distinctly different.

Mesosaurus To add credibility to his argument, Wegener documented cases of several fossil organisms found on different landmasses despite the unlikely pos-



sibility that their living forms could have crossed the vast ocean presently separating them (Figure 2.4). A classic example is Mesosaurus, a small aquatic freshwater reptile whose fossil remains are limited to black shales of the Permian period (about 260 million years ago) in eastern South America and southwestern Africa. If Mesosaurus had been able to make the long journey across the South Atlantic, its remains would likely be more widely distributed. As this is not the case, Wegener asserted that South America and Africa must have been joined during that period of Earth history.

How did opponents of continental drift explain the existence of identical fossil organisms in places separated by thousands of kilometers of open ocean? Rafting, transoceanic land bridges (isthmian links), and island stepping stones were the most widely invoked explanations for these migrations (Figure 2.5). We know, for example, that during the Ice Age that ended about 8000 years ago, the lowering of sea level allowed mammals (including humans) to cross the narrow Bering Strait that separates Russia and Alaska. Was it possible that land bridges once connected Africa and South America but later subsided below sea level? Modern maps of the seafloor substantiate Wegener's contention that if land bridges of this magnitude once existed, their remnants would still lie below sea level.

Glossopteris Wegener also cited the distribution of the fossil "seed fern" Glossopteris as evidence for the existence of Pangaea (see Figure 2.4). This plant, identified by its tongue-shaped leaves and seeds that were too large to be carried by the wind, was known to be widely dispersed among Africa, Australia, India, and South America. Later, fossil remains of Glossopteris were also discovered in Antarctica.* Wegener also learned that these seed ferns and associated flora grew only in cool climates-similar to central Alaska. Therefore, he concluded that when these landmasses were joined, they were located much closer to the South Pole.

Evidence: Rock Types and Geologic Features

Anyone who has worked a jigsaw puzzle knows that its successful completion requires that you fit the pieces together while maintaining the continuity of the picture.

*In 1912 Captain Robert Scott and two companions froze to death lying beside 35 pounds (16 kilograms) of rock on their return from a failed attempt to be the first to reach the South Pole. These samples, collected on the moraines of Beardmore Glacier, contained fossil remains of Glossopteris.

Figure 2.4 Fossil evidence supporting continental drift

Fossils of identical organisms have been discovered in rocks of similar age in Australia, Africa, South America, Antarctica, and India—continents that are currently widely separated by ocean barriers. Wegener accounted for these occurrences by placing these continents in their pre-drift locations.

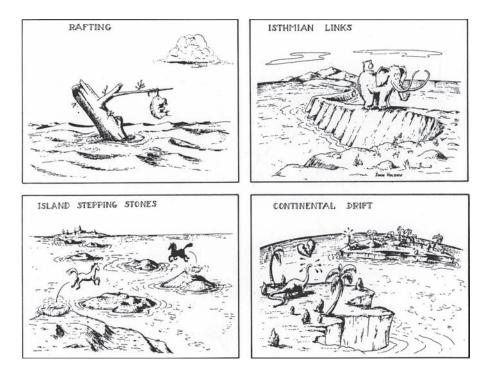
The "picture" that must match in the "continental drift puzzle" is one of rock types and geologic features such as mountain belts. If the continents were together, the rocks found in a particular region on one continent should closely match in age and type those found in adjacent positions on the once adjoining continent. Wegener found evidence of 2.2-billion-year-old igneous rocks in Brazil that closely resembled similarly aged rocks in Africa.

Similar evidence can be found in mountain belts that terminate at one coastline and reappear on landmasses across the ocean. For instance, the mountain belt that includes the Appalachians trends northeastward through the eastern United States and disappears off the coast of Newfoundland (Figure 2.6A). Mountains of comparable age and structure are found in the British Isles, western Africa, and Scandinavia. When these landmasses are positioned as they were about 200 million years ago, as shown in Figure 2.6B, the mountain chains form a nearly continuous belt.

Wegener described how the similarities in geologic features on both sides of the Atlantic linked these landmasses when he said, "It is just as if we were to refit the torn pieces of a newspaper by matching their edges and then check whether the lines of print run smoothly across. If they do, there is nothing left but to conclude that the pieces were in fact joined in this way."*

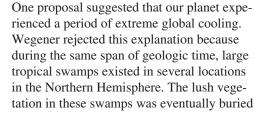
Evidence: Ancient Climates

Because Alfred Wegener was a student of world climates, he suspected that paleoclimatic (*paleo* = ancient, *climatic* = climate) data might also support the idea of mobile continents. His assertion was bolstered when he learned that



evidence for a glacial period that dated to the late Paleozoic had been discovered in southern Africa, South America, Australia, and India. This meant that about 300 million years ago, vast ice sheets covered extensive portions of the Southern Hemisphere as well as India (**Figure 2.7A**). Much of the land area that contains evidence of this period of Paleozoic glaciation presently lies within 30° of the equator in subtropical or tropical climates.

How could extensive ice sheets form near the equator?



South

America

Figure 2.5 How do land animals cross vast

oceans? These sketches illustrate various explanations for the occurrence of similar species on landmasses that are presently separated by vast oceans. (Used by permission of John Holden)



Africa

Figure 2.6 Matching mountain ranges across the North Atlantic



*Alfred Wegener, *The Origin of Continents and Oceans*, translated from the 4th revised German ed. of 1929 by J. Birman (London: Methuen, 1966).

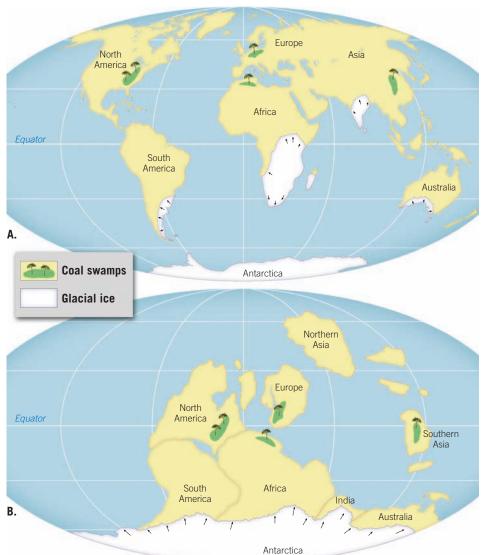


Figure 2.7 Paleoclimatic evidence for continental drift A. About 300 mil-

lion years ago, ice sheets covered extensive areas of the Southern Hemisphere and India. Arrows show the direction of ice movement that can be inferred from the pattern of glacial striations and grooves found in the bedrock. **B**. The continents restored to their pre-drift positions accounts for tropical coal swamps that existed in areas presently located in temperate climates. and converted to coal (Figure 2.7B). Today these deposits comprise major coal fields in the eastern United States and Northern Europe. Many of the fossils found in these coal-bearing rocks were produced by tree ferns that possessed large fronds—a feature consistent with warm, moist climates.** The existence of these large tropical swamps, Wegener argued, was inconsistent with the proposals that extreme global cooling caused glaciers to form in what are currently tropical areas.

Wegener suggested that a more plausible explanation for the late Paleozoic glaciation was provided by the supercontinent of Pangaea. In this configuration, the southern continents are joined together and located near the South Pole (see Figure 2.7B). This would account for the conditions necessary to generate extensive expanses of glacial ice over much of these landmasses. At the same time, this geography would place today's northern continents nearer the equator and account for the tropical swamps that generated the vast coal deposits.

How does a glacier develop in hot, arid central Australia? How do land animals migrate across wide expanses of the ocean? As compelling as this evidence may have been, 50 years passed before most of the scientific community accepted the concept of continental drift and the logical conclusions to which it led.

Concept Checks 2.2

- (1) What was the first line of evidence that led early investigators to suspect that the continents were once connected?
- Explain why the discovery of the fossil remains of Mesosaurus in both South America and Africa, but nowhere else, supports the continental drift hypothesis.
- (3) Early in the twentieth century, what was the prevailing view of how land animals migrated across vast expanses of open ocean?
- (4) How did Wegener account for the existence of glaciers in the southern landmasses at a time when areas in North America, Europe, and Asia supported lush tropical swamps?

**It is important to note that coal can form in a variety of climates, provided that large quantities of plant life are buried.

3 The Great Debate

Discuss the two main objections to the continental drift hypothesis.

Wegener's proposal did not attract much open criticism until 1924, when his book was translated into English, French, Spanish, and Russian. From that point until his death in 1930, the drift hypothesis encountered a great deal of hostile criticism. The respected American geologist R. T. Chamberlain stated, "Wegener's hypothesis in general is of the foot-loose type, in that it takes considerable liberty with our globe, and is less bound by restrictions or tied down by awkward, ugly facts than most of its rival theories."

Rejection of the Drift Hypothesis

One of the main objections to Wegener's hypothesis stemmed from his inability to identify a credible mechanism for continental drift. Wegener proposed that gravitational forces of the Moon and Sun that produce Earth's tides were also capable of gradually moving the continents across the globe. However, the prominent physicist Harold Jeffreys correctly countered that tidal forces strong enough to move Earth's continents would have resulted in halting our planet's rotation, which, of course, has not happened.

Wegener also incorrectly suggested that the larger and sturdier continents broke through thinner oceanic crust, much as ice breakers cut through ice. However, no evidence existed to suggest that the ocean floor was weak enough to permit passage of the continents without the continents being appreciably deformed in the process.

In 1930 Wegener made his fourth and final trip to the Greenland Ice Sheet (Figure 2.8). Although the primary focus of this expedition was to study this great ice cap and its climate, Wegener continued to test his continental drift hypothesis. While returning from Eismitte, an experimental station located in the center of Greenland, Wegener perished along with his Greenland companion. His intriguing idea, however, did not die.

Why was Wegener unable to overturn the established scientific views of his day? Foremost was the fact that, although the central theme of Wegener's drift hypothesis was correct, it contained some incorrect details. For example, continents do not break through the ocean floor, and tidal energy is much too weak to cause continents to be displaced. Moreover, in order for any comprehensive scientific theory to gain wide acceptance, it must withstand critical testing from all areas of science. Despite Wegener's great contribution to our understanding of Earth, not *all* of the evidence supported the continental drift hypothesis as he had proposed it.

Although many of Wegener's contemporaries opposed his views, even to the point of open ridicule, some considered his ideas plausible. For those geologists who continued the search, the exciting concept of continents adrift held their interest. Others viewed continental drift as a solution to previously unexplainable observations such

 Alfred Wegener shown waiting out the

 1912-1913 Arctic winter during an

 expedition to Greenland, where he

 made a 1200-kilometer traverse across

 the widest part of the island's ice sheet.

as the cause of earthquakes. Nevertheless, most of the scientific community, particularly in North America, either categorically rejected continental drift or treated it with considerable skepticism.

Concept Checks 2.3

- (1) What analogy did Wegener use to describe how the continents move though the ocean floor?
- What two aspects of Wegener's continental drift hypothesis were objectionable to most Earth scientists?

2.4 The Theory of Plate Tectonics

List the major differences between Earth's lithosphere and asthenosphere and explain the importance of each in the plate tectonics theory.

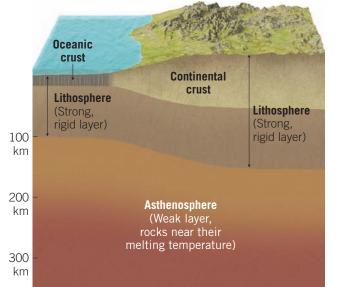
Following World War II, oceanographers equipped with new marine tools and ample funding from the U.S. Office of Naval Research embarked on an unprecedented period of oceanographic exploration. Over the next two decades, a much better picture of large expanses of the seafloor slowly and painstakingly began to emerge. From this work came the discovery of a global **oceanic ridge system** that winds through all the major oceans in a manner similar to the seams on a baseball.

In other parts of the ocean, more new discoveries were being made. Studies conducted in the western Pacific demonstrated that earthquakes were occurring at great depths beneath deep-ocean trenches. Of equal importance was the fact that dredging of the seafloor did not bring up any oceanic crust that was older than 180 million years. Further, sediment accumulations in the deep-ocean basins were found to be thin, not the thousands of meters that were predicted. By 1968 these developments, among others, had led to the unfolding of a far more encompassing theory than continental drift, known as the **theory of plate tectonics** (*tekto* = to build). Figure 2.8 Alfred Wegener during an expedition to Greenland (Photo courtesy of Archive of Alfred Wegener Institute)

Did You Know?

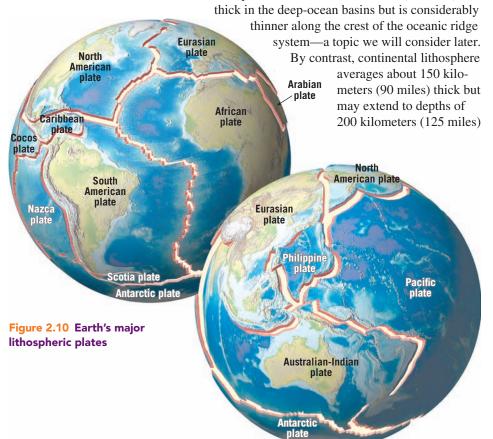
Alfred Wegener, best known for his continental drift hypothesis, also wrote numerous scientific papers on weather and climate. Because of his interest in meteorology, Wegener made four extended trips to the Greenland Ice Sheet in order to study its harsh winter weather. In November 1930, while making a month-long trek across the ice sheet, Wegener and a companion perished. SmartFigure 2.9 Rigid lithosphere overlies the weak asthenosphere





Rigid Lithosphere Overlies Weak Asthenosphere

According to the plate tectonics model, the crust and the uppermost, and therefore coolest, part of the mantle constitute Earth's strong outer layer, known as the **lithosphere** (*lithos* = stone, *sphere* = ball). The lithosphere varies in both thickness and density, depending on whether it is oceanic lithosphere or continental lithosphere (**Figure 2.9**). Oceanic lithosphere is about 100 kilometers (60 miles)



or more beneath the stable interiors of the continents. Further, the composition of both the oceanic and continental crusts affects their respective densities. Oceanic crust is composed of rocks that have a mafic (basaltic) composition, and therefore oceanic lithosphere has a greater density than continental lithosphere. Continental crust is composed largely of less dense felsic (granitic) rocks, making continental lithosphere less dense than its oceanic counterpart.

The **asthenosphere** (*asthenos* = weak, *sphere* = ball) is a hotter, weaker region in the mantle that lies below the lithosphere (see Figure 2.9). The temperatures and pressures in the upper asthenosphere (100 to 200 kilometers [60 to 175 miles] in depth) are such that rocks at this depth are very near their melting temperatures and, hence, respond to forces by *flowing*, similarly to the way a thick liquid would flow. By contrast, the relatively cool and rigid lithosphere tends to respond to forces acting on it by *bending or breaking but not flowing*. Because of these differences, Earth's rigid outer shell is effectively detached from the asthenosphere, which allows these layers to move independently.

Earth's Major Plates

The lithosphere is broken into about two dozen segments of irregular size and shape called lithospheric plates, or simply **plates**, that are in constant motion with respect to one another (Figure 2.10). Seven major lithospheric plates are recognized and account for 94 percent of Earth's surface area: the North American, South American, Pacific, African, Eurasian, Australian–Indian, and Antarctic plates. The largest is the Pacific plate, which encompasses a significant portion of the Pacific basin. Each of the six other large plates includes an entire continent plus a significant amount of ocean floor. Notice in Figure 2.11 that the South American plate encompasses almost all of South America and about one-half of the floor of the South Atlantic. This is a major departure from Wegener's continental drift hypothesis, which proposed that the continents move through the ocean floor, not with it. Note also that none of the plates are defined entirely by the margins of a single continent.

Intermediate-sized plates include the *Caribbean*, *Nazca*, *Philippine*, *Arabian*, *Cocos*, *Scotia*, and *Juan de Fuca plates*. These plates, with the exception of the Arabian plate, are composed mostly of oceanic lithosphere. In addition, several smaller plates (*microplates*) have been identified but are not shown in Figure 2.11.

Plate Boundaries

One of the main tenets of the plate tectonics theory is that plates move as somewhat rigid units relative to all other plates. As plates move, the distance between two locations on different plates, such as New York and London, gradually changes, whereas the distance between sites on the same plate—New York and Denver, for example remains relatively constant. However, parts of some plates are comparatively "soft," such as southern China, which is literally being squeezed as the Indian subcontinent rams into Asia proper.

Because plates are in constant motion relative to each other, most major interactions among them (and, therefore, most deformation) occur along their *boundaries*. In fact, plate boundaries were first established by plotting the locations of earthquakes and volcanoes. Plates are bounded by three distinct types of boundaries, which are differentiated by the type of movement they exhibit. These boundaries are depicted in Figure 2.11 and are briefly described here:

- 1. Divergent plate boundaries (*constructive margins*) where two plates move apart, resulting in upwelling of hot material from the mantle to create new seafloor (Figure 2.11A).
- 2. Convergent plate boundaries (*destructive margins*) where two plates move together, resulting in oceanic lithosphere descending beneath an overriding plate, eventually to be reabsorbed into the mantle or possibly in the collision of two continental blocks to create a mountain belt (Figure 2.11B).

 Transform plate boundaries (conservative margins)—where two plates grind past each other without producing or destroying lithosphere (Figure 2.11C).

Divergent and convergent plate boundaries each account for about 40 percent of all plate boundaries. Transform plate boundaries account for the remaining 20 percent. In the following sections we will summarize the nature of the three types of plate boundaries.

Concept Checks 2.4

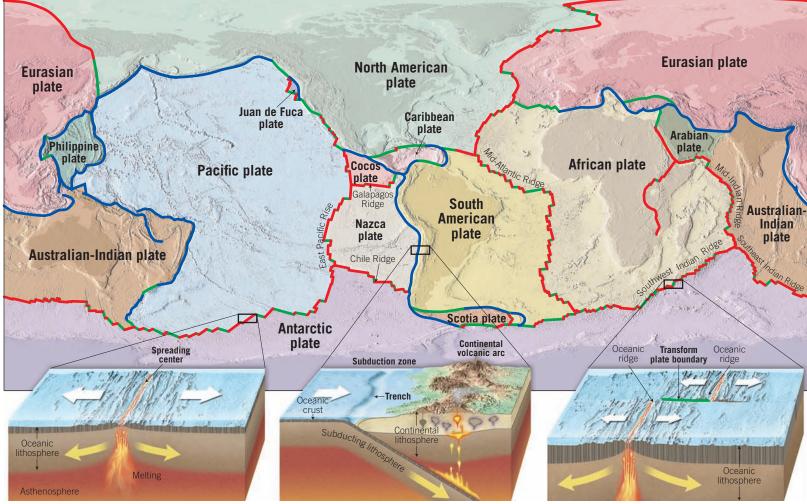
- (1) The global extent of what major ocean floor feature did oceanographers discover after World War II?
- (2) Compare and contrast the lithosphere and the asthenosphere.
- 3 List the seven largest lithospheric plates.
- (4) List the three types of plate boundaries and describe the relative motion at each of them.

Did You Know?

A group of scientists proposed an interesting although incorrect explanation for the cause of continental drift. Their proposal suggested that early in Earth's history, our planet was only about half its current diameter and completely covered by continental crust. Through time Earth expanded, causing the continents to split into their current configurations, while new seafloor "filled in" the spaces as they drifted apart.

Figure 2.11 Divergent, convergent, and transform plate boundaries

(Based on W. B. Hamilton, U.S. Geological Survey)



A. Divergent plate boundary -

B. Convergent plate boundary -

C. Transform plate boundary —